

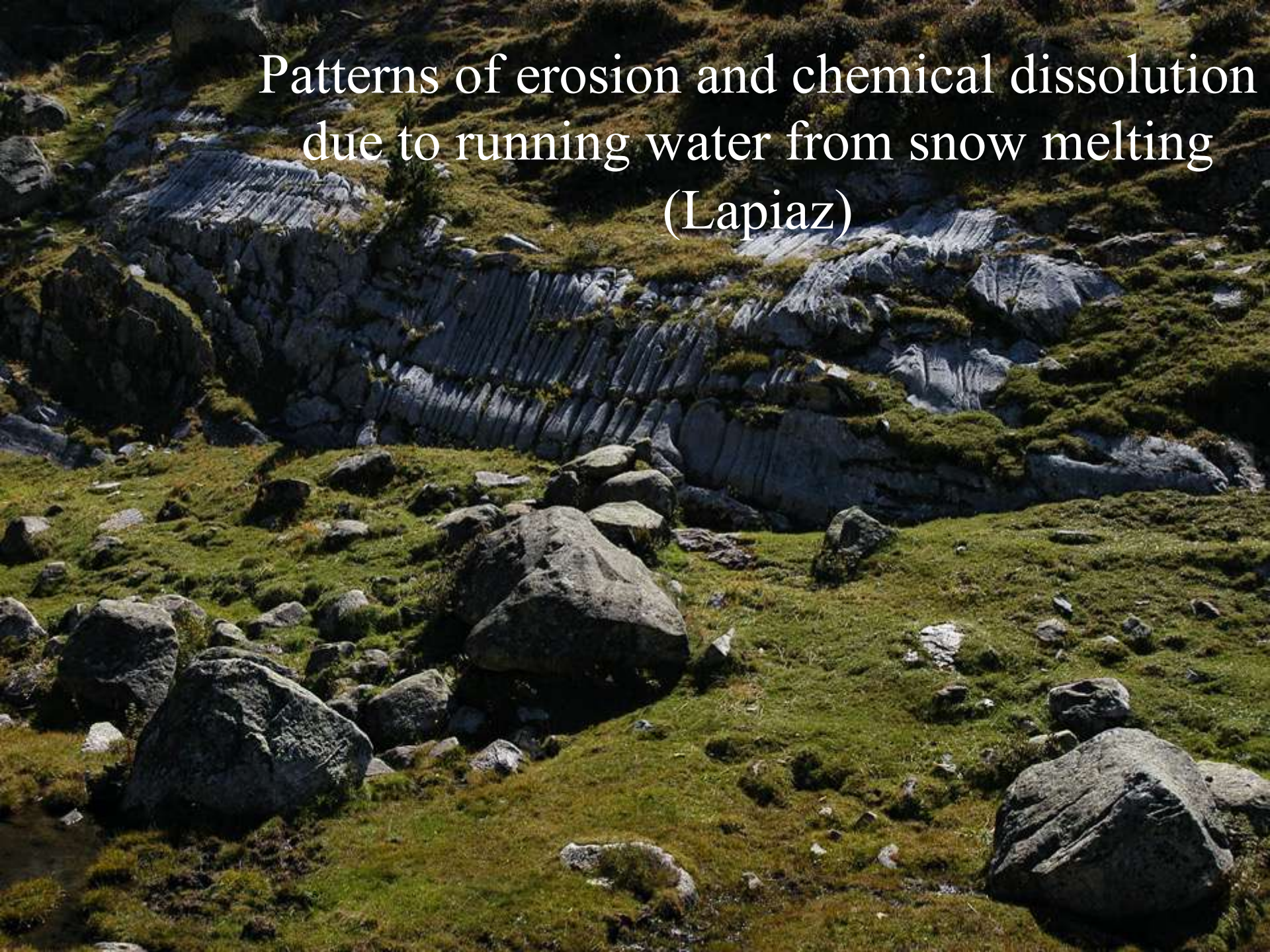
A scenic view of a mountain valley. In the foreground, there is a grassy slope leading down to a dark blue lake. The lake is surrounded by steep, rocky slopes that rise sharply on both sides. The sky is a clear, deep blue. The overall scene is a high-altitude mountain landscape.

Patterns in Geomorphology

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Patterns of erosion and chemical dissolution
due to running water from snow melting
(Lapiaz)





Patterns of erosion and chemical dissolution



Erosion patterns in badlands (Tzin valley, Israel)



Patterns of erosion and deposition:
Meandering rivers
(Ucayali river, Peru)

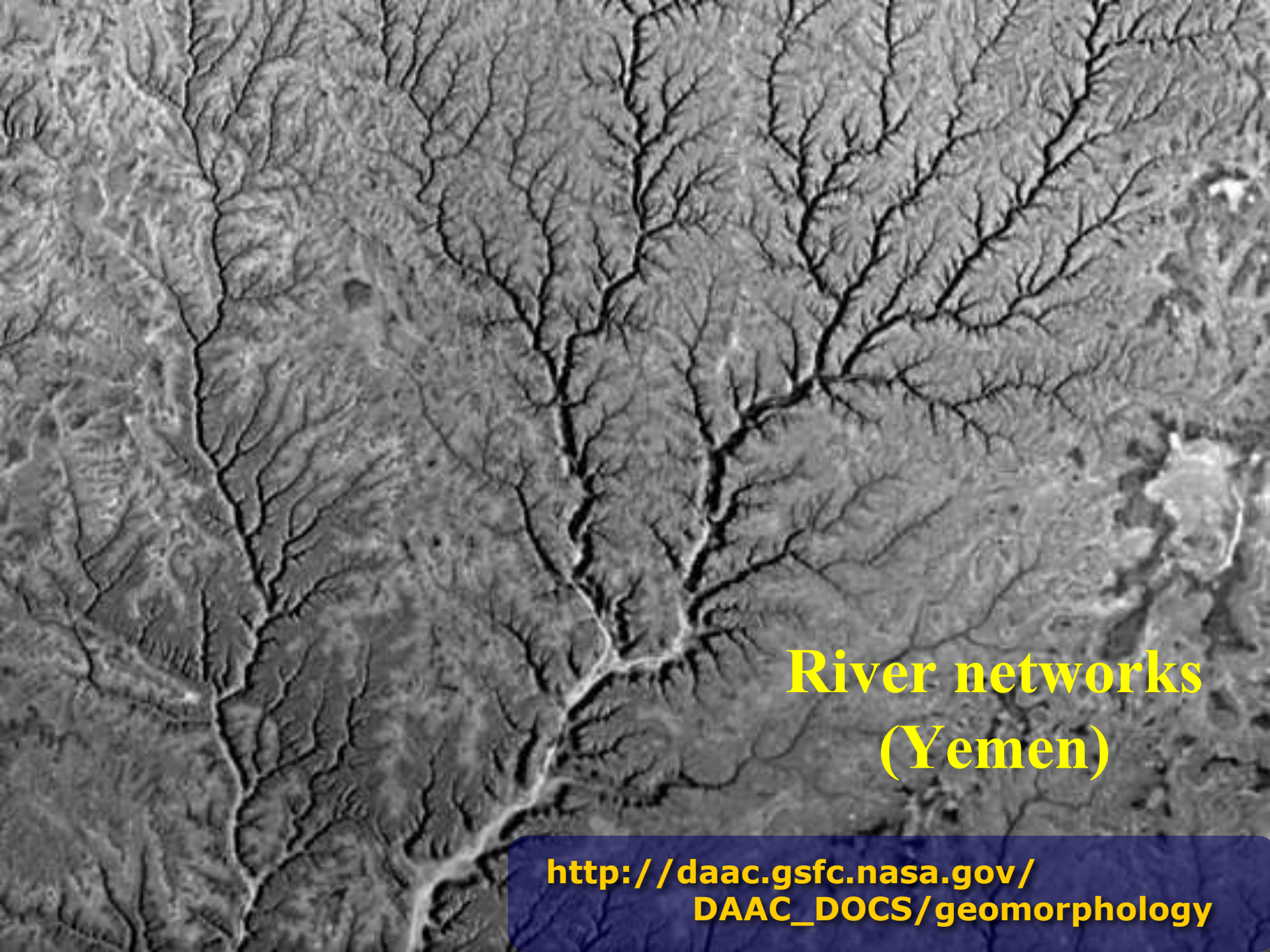


[http://daac.gsfc.nasa.gov/
DAAC_DOCS/geomorphology](http://daac.gsfc.nasa.gov/DAAC_DOCS/geomorphology)



Patterns of erosion
and deposition:
Braided rivers





River networks (Yemen)

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Coastal patterns (coast of Carolina, USA)



Courtesy of A. Brad Murray



Aeolian patterns
Great Sand Dunes National Monument, Colorado,
photo by Bob Bauer

Aeolian bedforms
in deserts and sandy beaches:

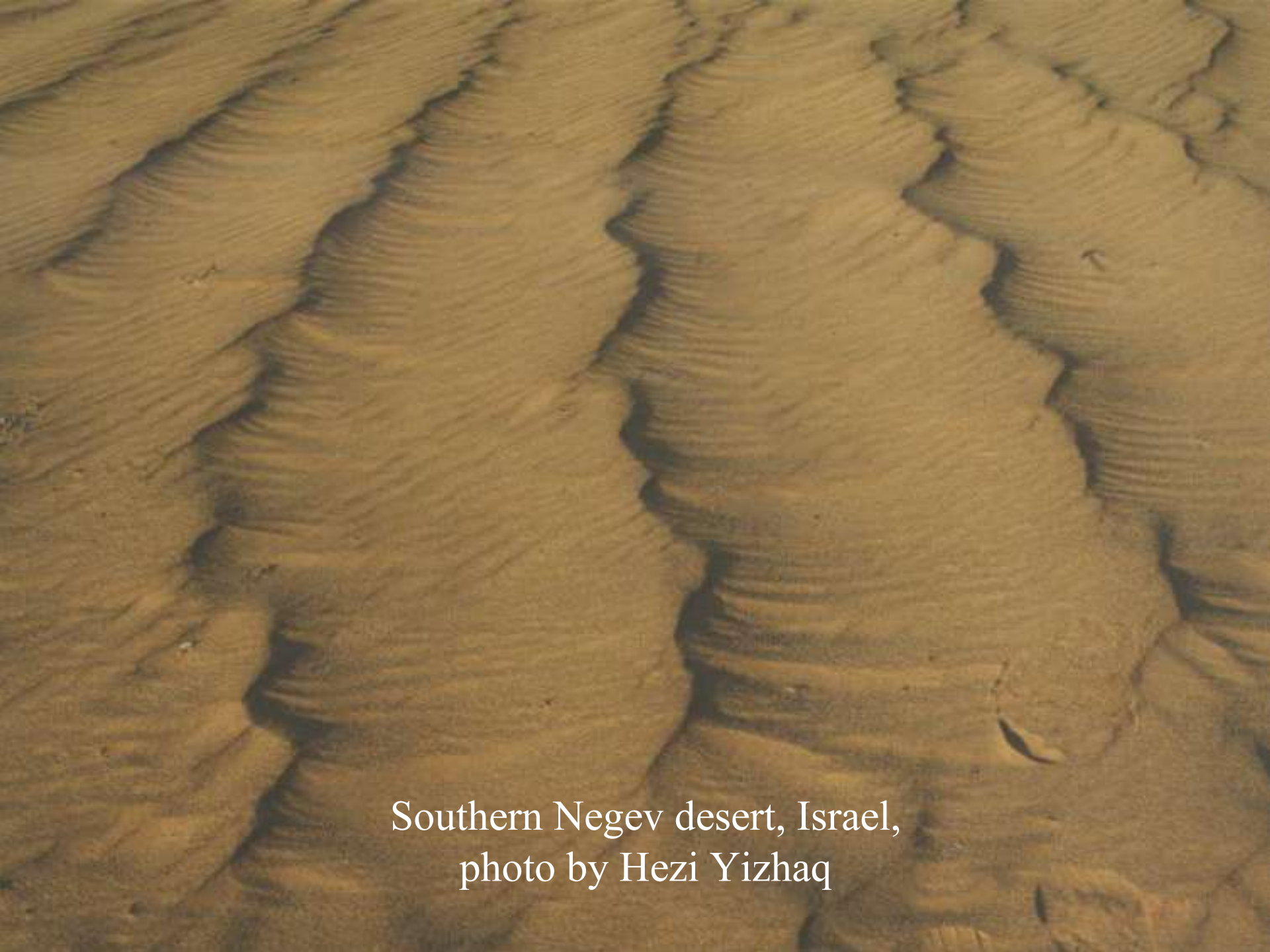
Ripples
(length of cm, amplitude of mm)

Megaripples
(length of meters, amplitude of cm)

Dunes
(length of tens or hundreds of meters,
amplitude of (tens of) meters)



Wadi Rum desert, Jordan



Southern Negev desert, Israel,
photo by Hezi Yizhaq

Types of (non-vegetated) dunes:

Barchans

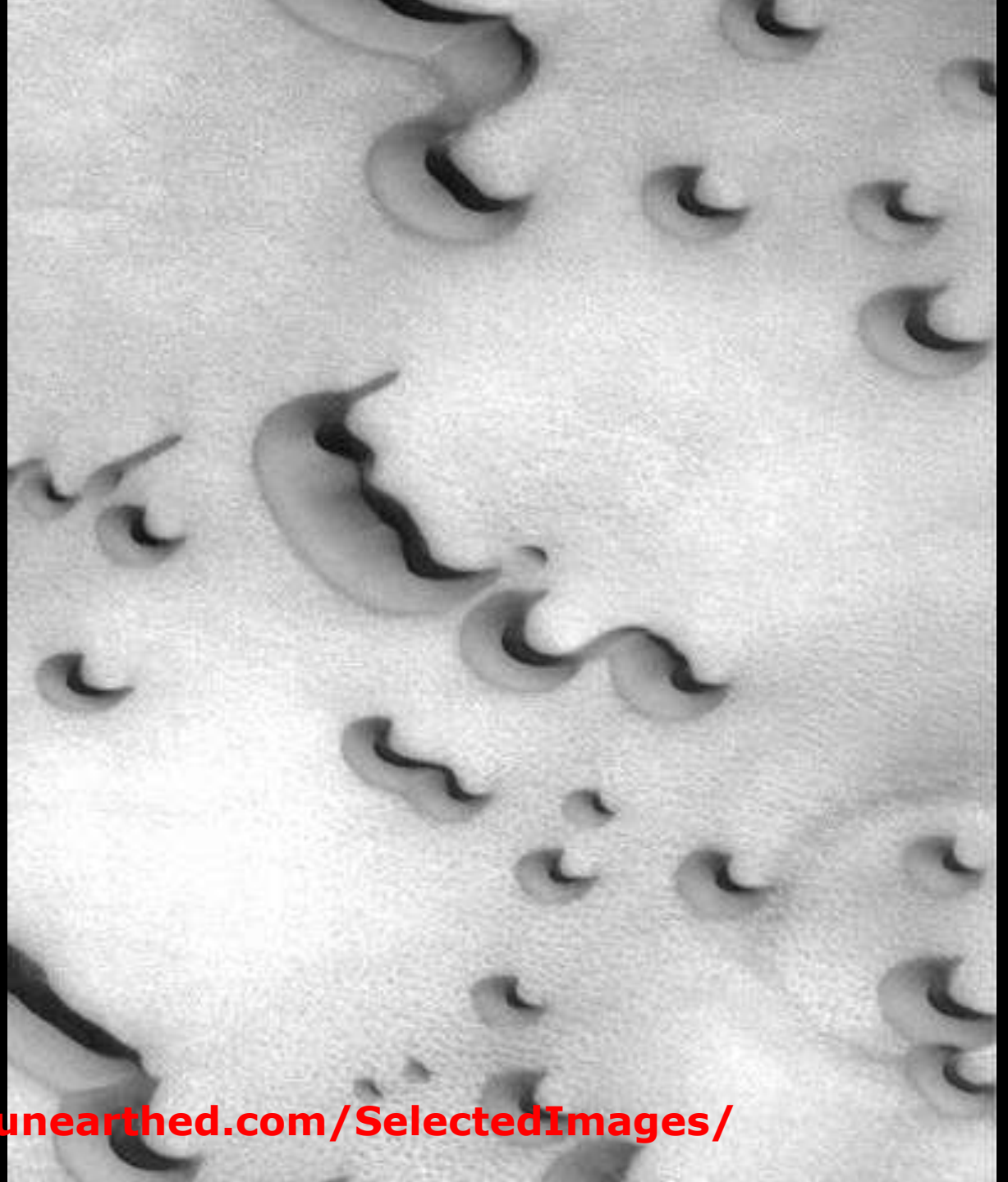
Transverse

Seif (linear) dunes

Star dunes



Barchan dunes
on Mars



<http://www.marsunearthed.com/SelectedImages/>

Namib desert



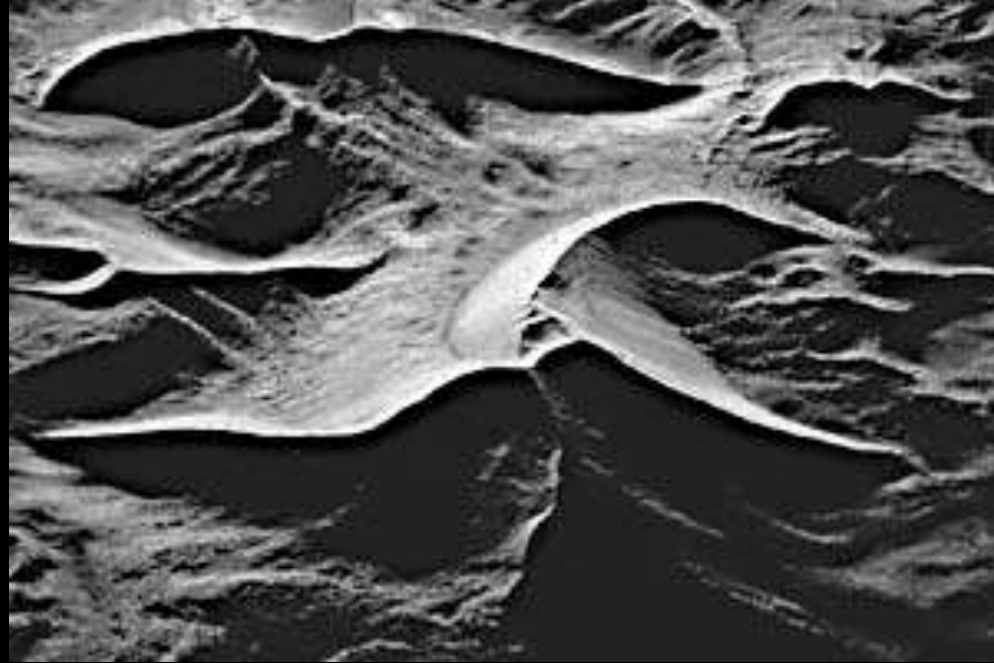
[http://daac.gsfc.nasa.gov/
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Linear dunes, Mauritania



[http://daac.gsfc.nasa.gov/
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Namib sand sea



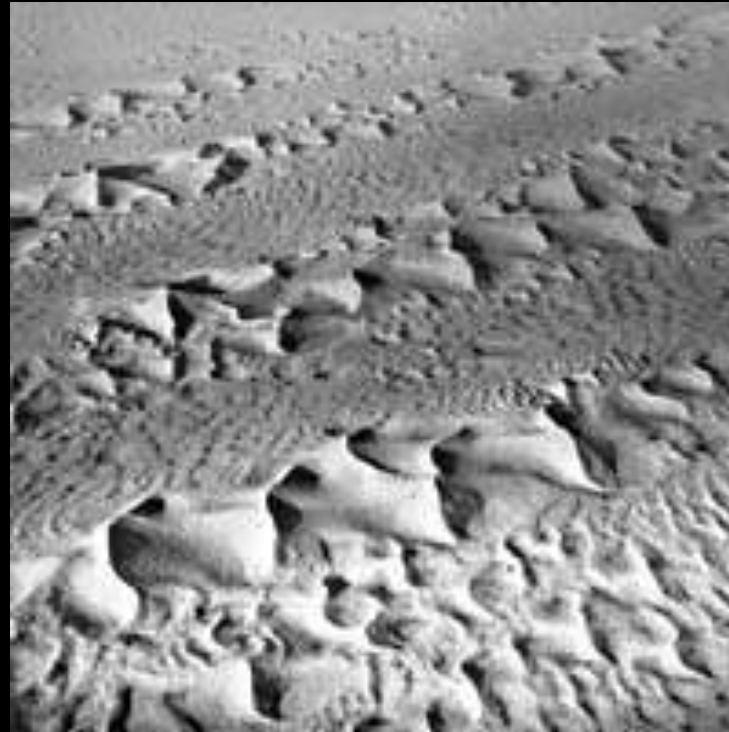
[http://daac.gsfc.nasa.gov/
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Complex dune structures (Saudi Arabia)



[http://daac.gsfc.nasa.gov/
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Complex dune structures (Gran Desierto))

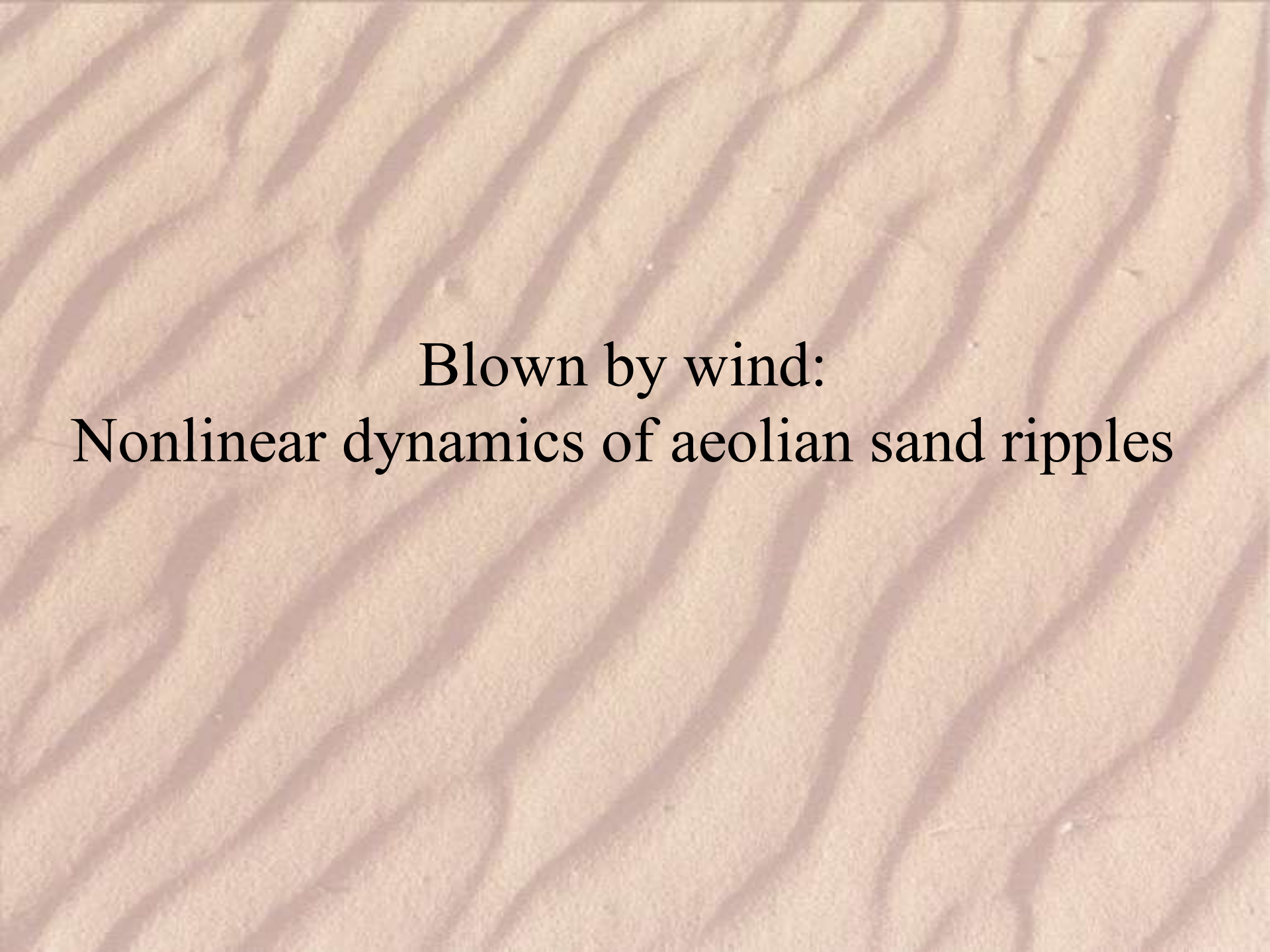


[http://daac.gsfc.nasa.gov/
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Superposed aeolian bedforms



[http://daac.gsfc.nasa.gov/
DAAC_DOCS/geomorphology](http://daac.gsfc.nasa.gov/DAAC_DOCS/geomorphology)

The background of the slide is a close-up photograph of sand ripples. The ripples are formed by wind-blown sand, creating a series of parallel, wavy ridges and troughs. The color of the sand is a light, warm beige. The lighting is soft, highlighting the texture and the undulating nature of the ripples.

Blown by wind:
Nonlinear dynamics of aeolian sand ripples

Properties of aeolian ripples:

Ripple index about 15-20

Almost 1D bedforms:

crests are perpendicular to the wind, with defects

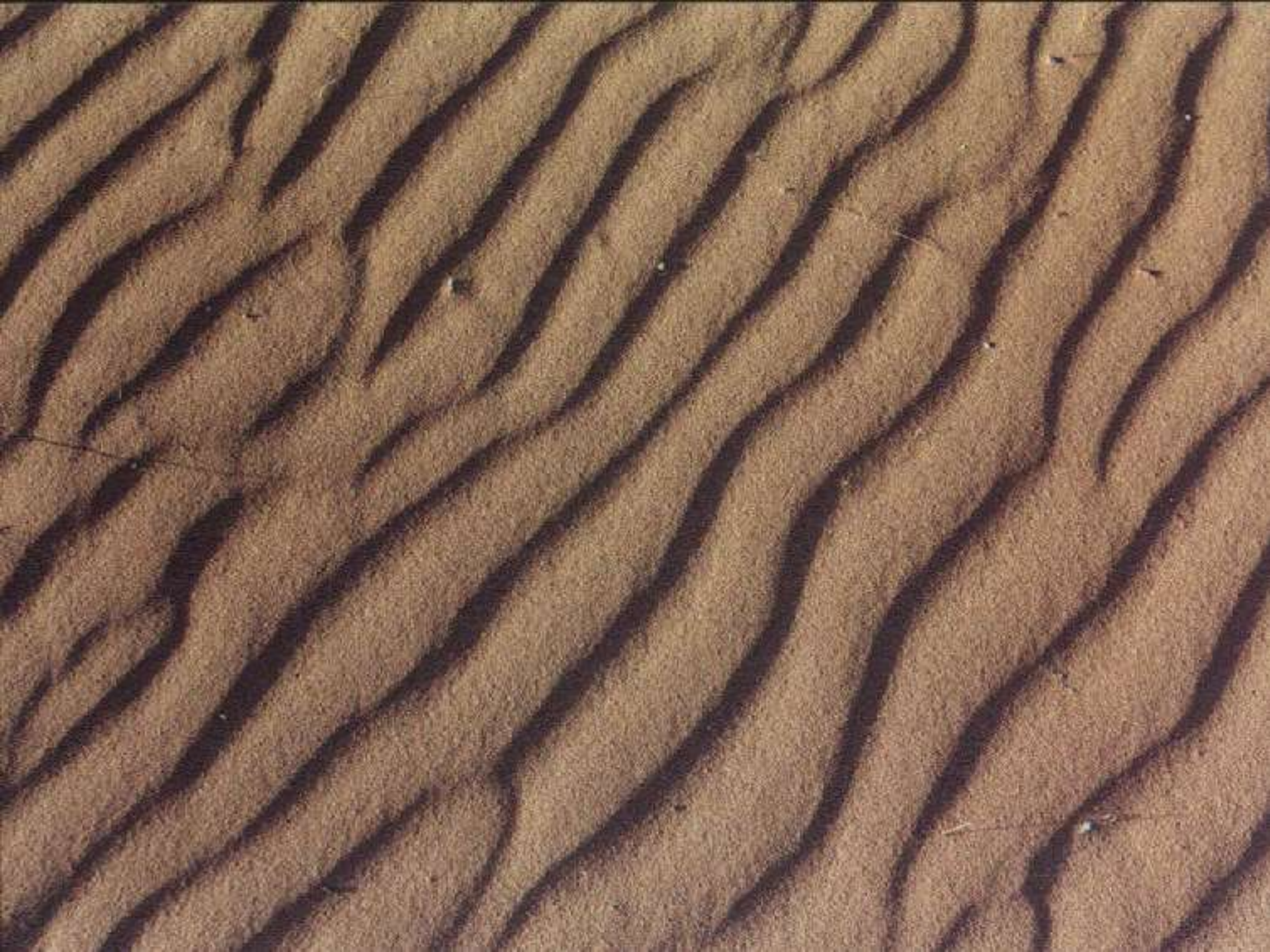
Slight asymmetry between lee and stoss slopes
(no slip face)

Need of a wind intensity threshold
to have ripple formation

Rapid response to wind

The ripple pattern coarsens with time
and it slowly moves downstream

Grain sorting



A short history of ripple studies:

R.A. Bagnold, *The Physics of Blown Sand and Desert Dunes*, 1941

R. Cooke, A. Warren, A. Goudie, *Desert Geomorphology*, 1993

N. Lancaster, *Geomorphology of desert dunes*, 1995

R.P. Sharp, *J. of Geology*, 1963

M. Seppala and K. Lindé, *Geografiska Annaler*, 1978

B.B. Willetts and M.A. Rice, 1983-1989

J.E. Ungar and P.K. Haff, *Sedimentology*, 1987

R.S. Anderson, *Sedimentology*, 1987

R.S. Anderson, *Earth Sci. Rev.*, 1990

B.T. Werner and D.T. Gillespie, *PRL*, 1993

W. Landry and B.T. Werner, *Physica D*, 1994

H. Nishimori and N. Ouchi, *PRL*, 1993

R.B. Hoyle and A.W. Woods, *PRE*, 1997

L. Prigozhin, *PRE*, 1999

O. Terzidis, P. Claudin and J.-P. Bouchaud, *Eur. Phys. J. B*, 1998

A. Valance and F. Rioual, *Eur. Phys. J. B*, 1999

Z. Csahok and C. Misbah, *Eur. Phys. J. E*, 2000

Z. Csahok, C. Misbah, F. Rioual and A. Valance, *cond-mat*, 2000

H. Yizhaq, N.J. Balmforth, A. Provenzale, *Physica D*, 2004

H. Yizhaq et al. EPSI, 2019 (megaripples)

Mechanism of ripple formation:
aeolian ripples form due to
the instability of a flat sand bed
exposed to strong wind

When the wind starts to blow,
sand grains are lifted into the air.
These grains are accelerated
by the wind and fall down,
hit the surface,
and eject other grains.

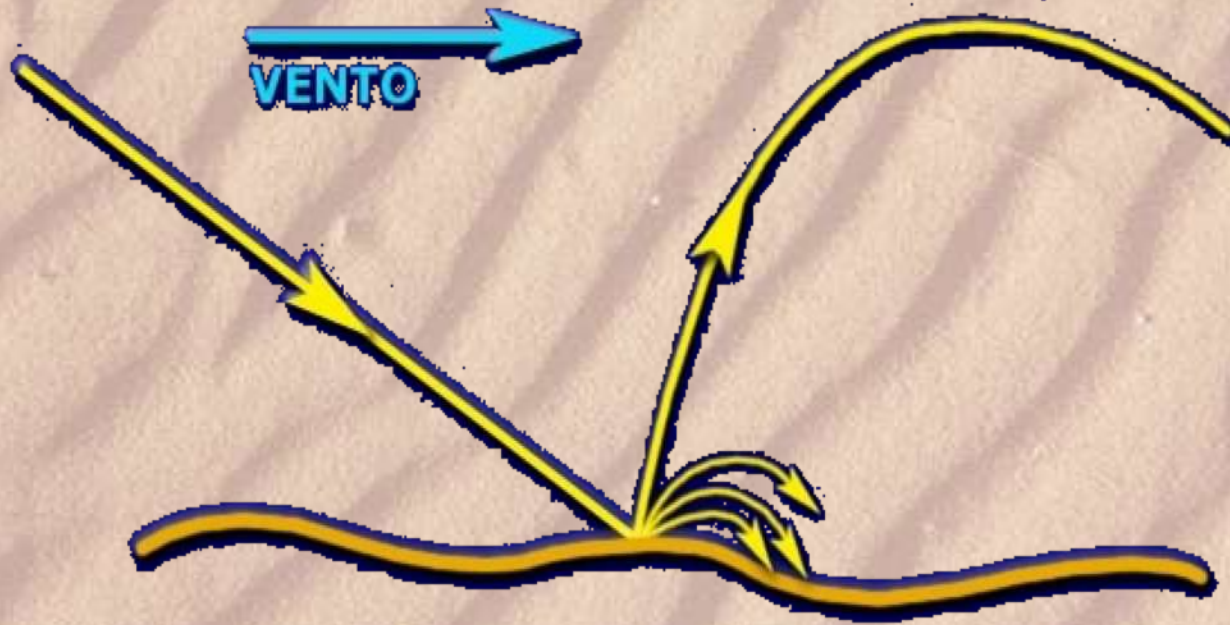
The rebounding (saltating) grains are then accelerated by the wind, and a cascade process ensues.

An entire population of saltating grains emerges.

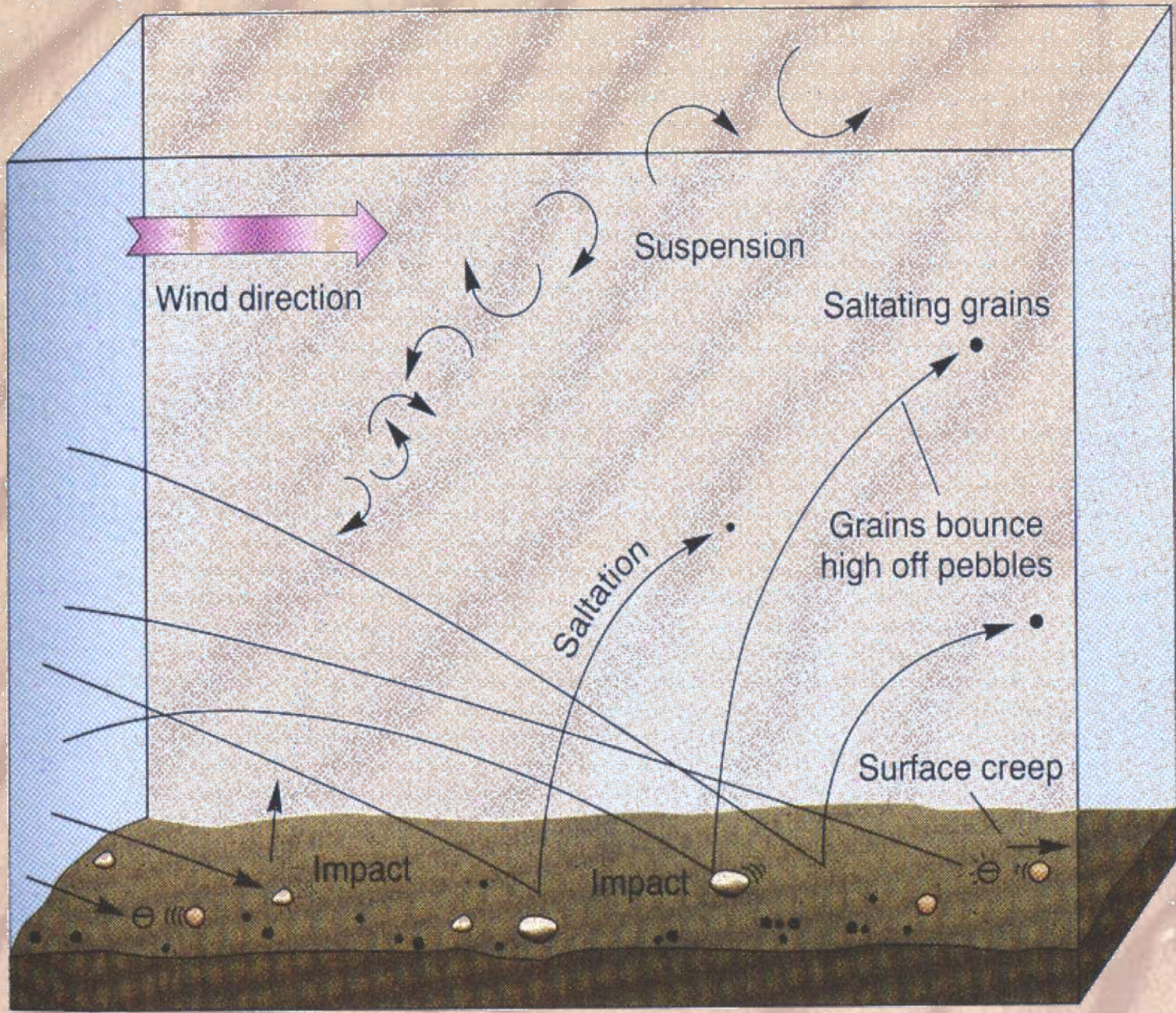
The height of the saltation layer can be about one meter in strong winds

Hypotheses of ripple formation (Cooke et al, 1993):

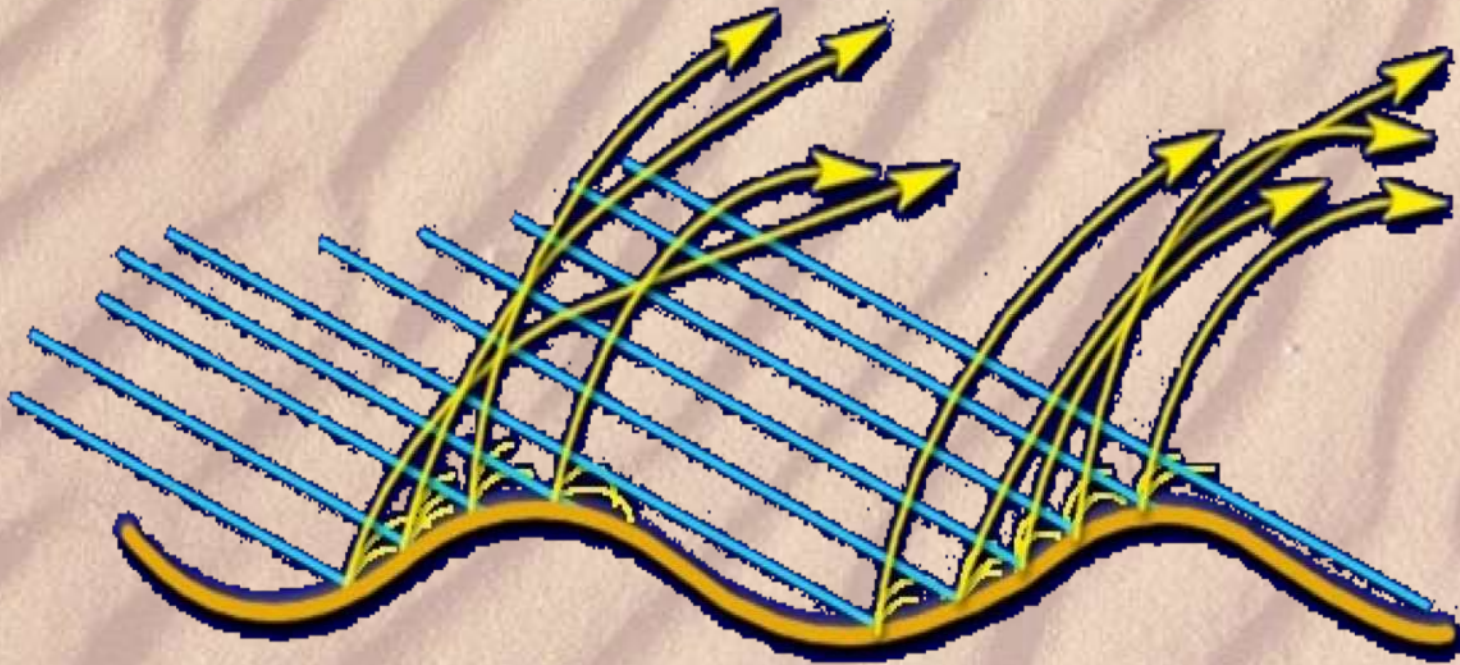
1. A rhythmic barrage of saltating grains (Bagnold)
2. The wave hypothesis:
 - a) The bed as a fluid
 - b) The saltation curtain as a fluid
 - c) Wave-like instabilities in the boundary layer
 - d) Secondary motions in the lee of transverse ripples
3. The role of reptating grains (Anderson 1987)



Saltation with typical jump length L
Reptation with typical jump length $a \ll L$
 L is about 1 m and a is about 1 cm

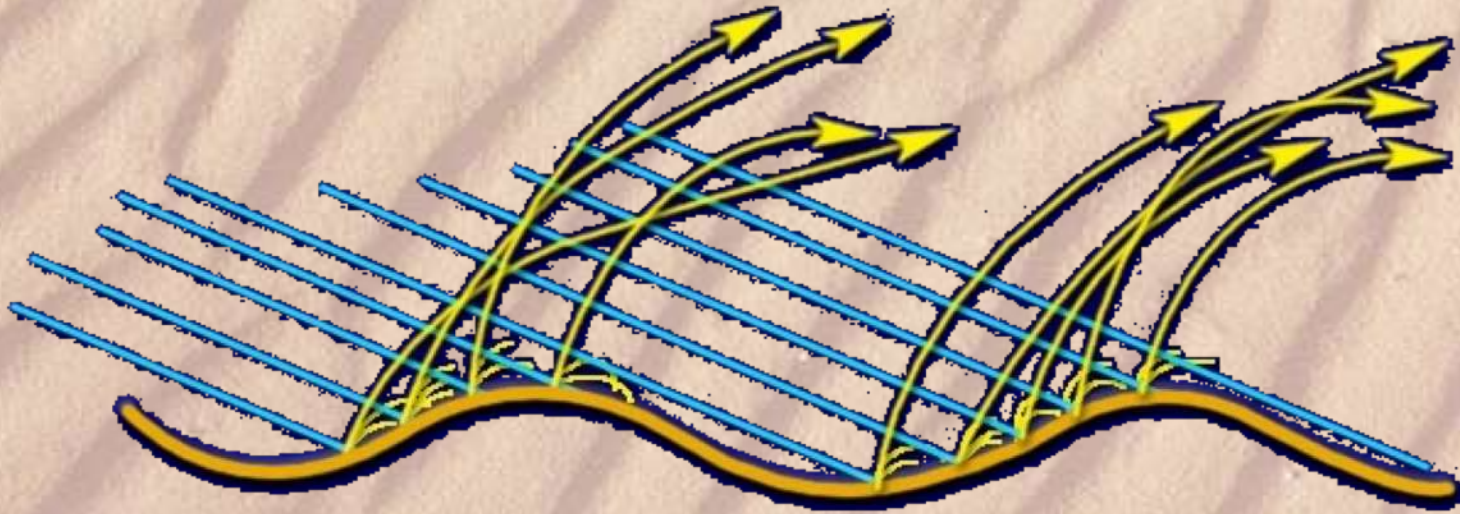


Upon impact,
the energy of a saltating grain goes as:
80% to one (on average) rebounding grain,
10% to a few reptating grains,
10% to the deformation of the bed



Depth of the saltation layer: up to about 1 m

Depth of the reptation layer: a few mm



We can idealize the problem in terms of a sand surface bombarded by a continuous flux of saltating grains that hit the surface at constant (small) angle $\phi = 8-12^\circ$

The saltating grains drive the system. The important dynamics is contained in the behavior of the reptating grains

Conservation of sand

$$(1 - n)\rho \frac{\partial \zeta}{\partial t} = -\nabla \cdot \mathbf{Q}$$

- n porosity of the bed (about 0.35)
- ρ density of sand
- ζ elevation of the sand surface
- \mathbf{Q} flux of sand grains

The flux of sand:

$$Q = Q_s + Q_r$$

- Q_s flux of saltating grains
- Q_r flux of reptating grains

The flux of saltating grains
is assumed to be constant for
aeolian ripples

$$\nabla \cdot \mathbf{Q}_s = 0$$

All the dynamics is contained in the
variability of the reptation flux

NB: This is untenable for megaripples and dunes

There is no feedback of the
aeolian bedforms
on the wind
and on the flux of saltating grains

NB: This is untenable for dunes

1D case

if all the grains had the same reptation length a

$$Q_r^{bare}(x, t) = m \int_{x-a}^x N_{ej}(x') dx'$$

- m mass of a sand grain
- N_{ej} number of ejected grains
- N_r average number of reptating grains ejected by one saltating grain
- $N_{ej} = N_r N_{im}$ where N_{im} is the number of impacting grains

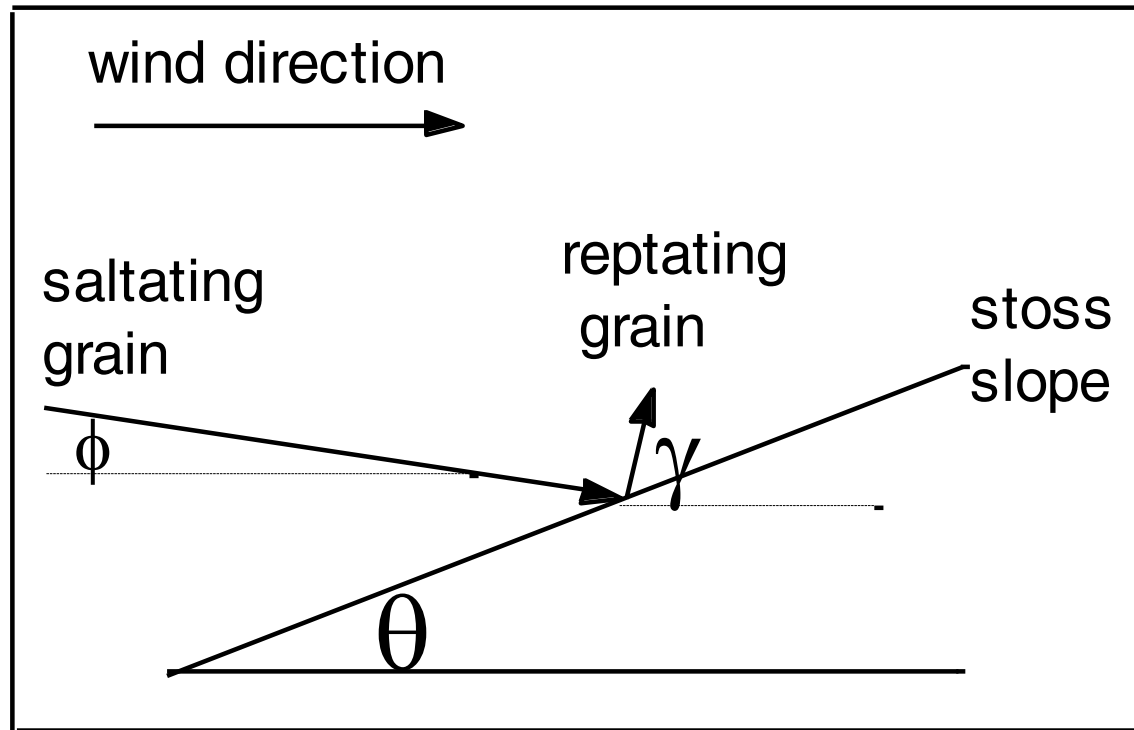
For a distribution of reptation lengths

$$Q_r^{bare}(x, t) = m \int_{-\infty}^{\infty} d\alpha p(\alpha) \int_{x-\alpha}^x N_{ej}(x') dx'$$

- m mass of a sand grain
- N_{ej} number of ejected grains
- N_r average number of reptating grains ejected by one saltating grain
- $N_{ej} = N_r N_{im}$ where N_{im} is the number of impacting grains
- $p(\alpha)$ distribution of reptation lengths

$$\int \alpha p(\alpha) d\alpha = a$$

Important angles



Number density of impacting grains

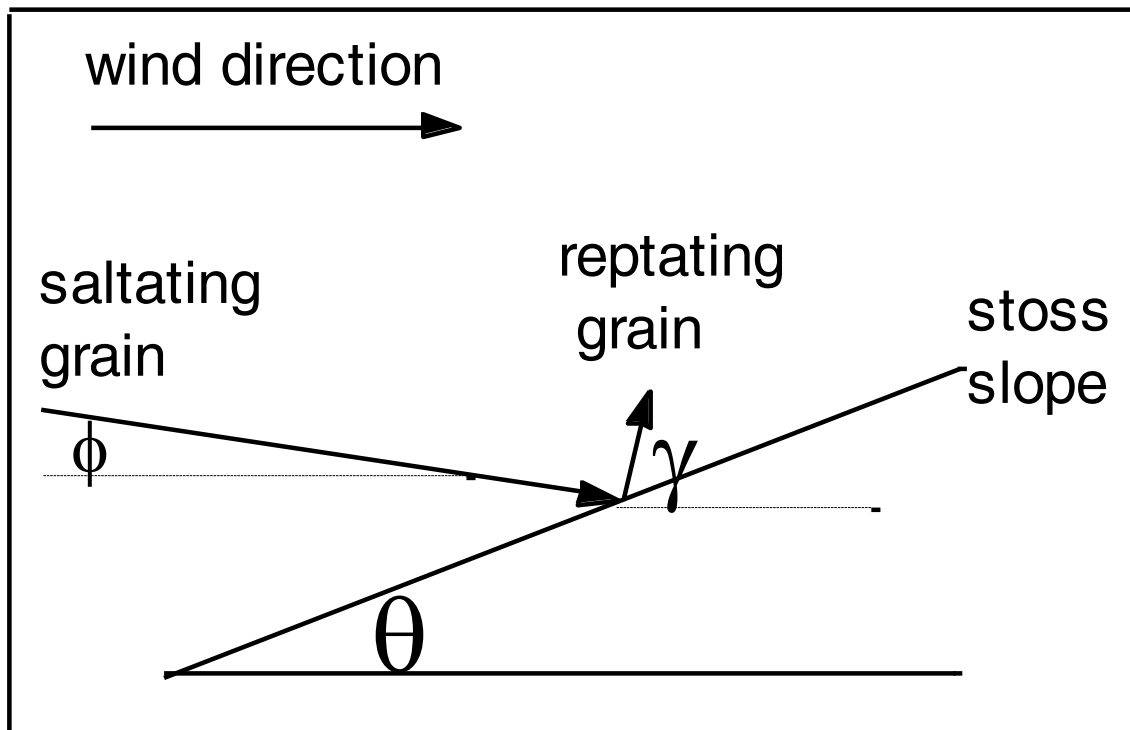
$$N_{im}(x) = N_{im}^0 \left(1 + \frac{\tan \theta}{\tan \phi} \right) \cos \theta$$
$$= N_{im}^0 \cot \phi \frac{\tan \phi + \zeta_x}{\sqrt{1 + \zeta_x^2}}$$

N_{im}^0 number density of impacting grains
on a horizontal surface (about $10^7 \text{ m}^{-2} \text{ s}^{-1}$)

θ inclination of the surface

$$\tan \theta = \zeta_x$$

$$a = 2 \frac{V^2}{g} \frac{\sin(\gamma - \theta) \cos \gamma}{\cos \theta} = a_{hor} (1 - \cot \gamma \tan \theta)$$



Shadowing:
the flux of reptating grains
becomes

$$N_{im}(x) = N_{im}^0 \cot \phi \operatorname{Max} \left\{ \frac{\tan \phi + \zeta_x}{\sqrt{1 + \zeta_x^2}}, 0 \right\}$$

The full integral model:

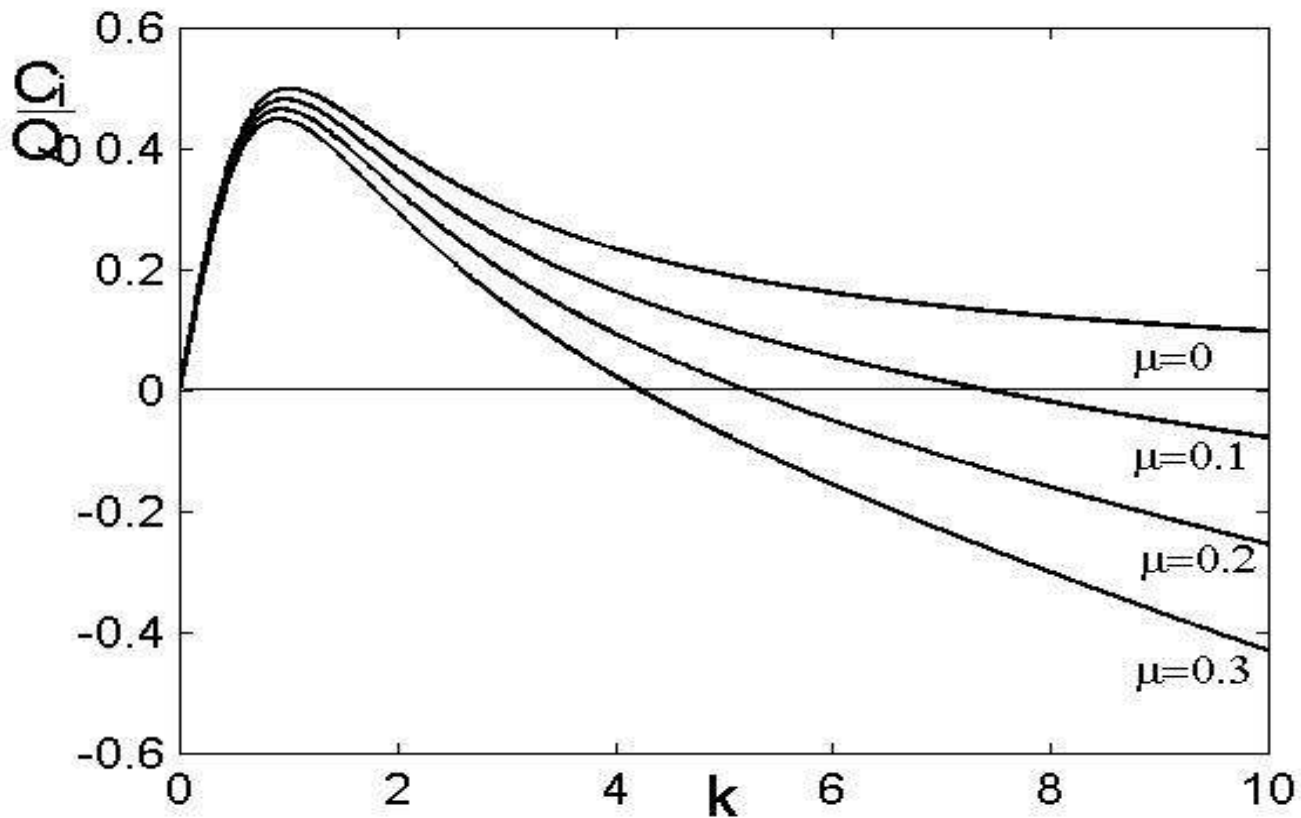
$$\frac{\partial \zeta}{\partial t} = -Q_0 \frac{\partial}{\partial x} \left[(1 - \mu \zeta_x) \int_{-\infty}^{\infty} d\alpha p(\alpha) \int_{x-\alpha}^x F(x') dx' \right]$$

$$F(x) = \text{Max} \left\{ \frac{\tan \phi + \zeta_x}{\sqrt{1 + \zeta_x^2}}, 0 \right\}$$

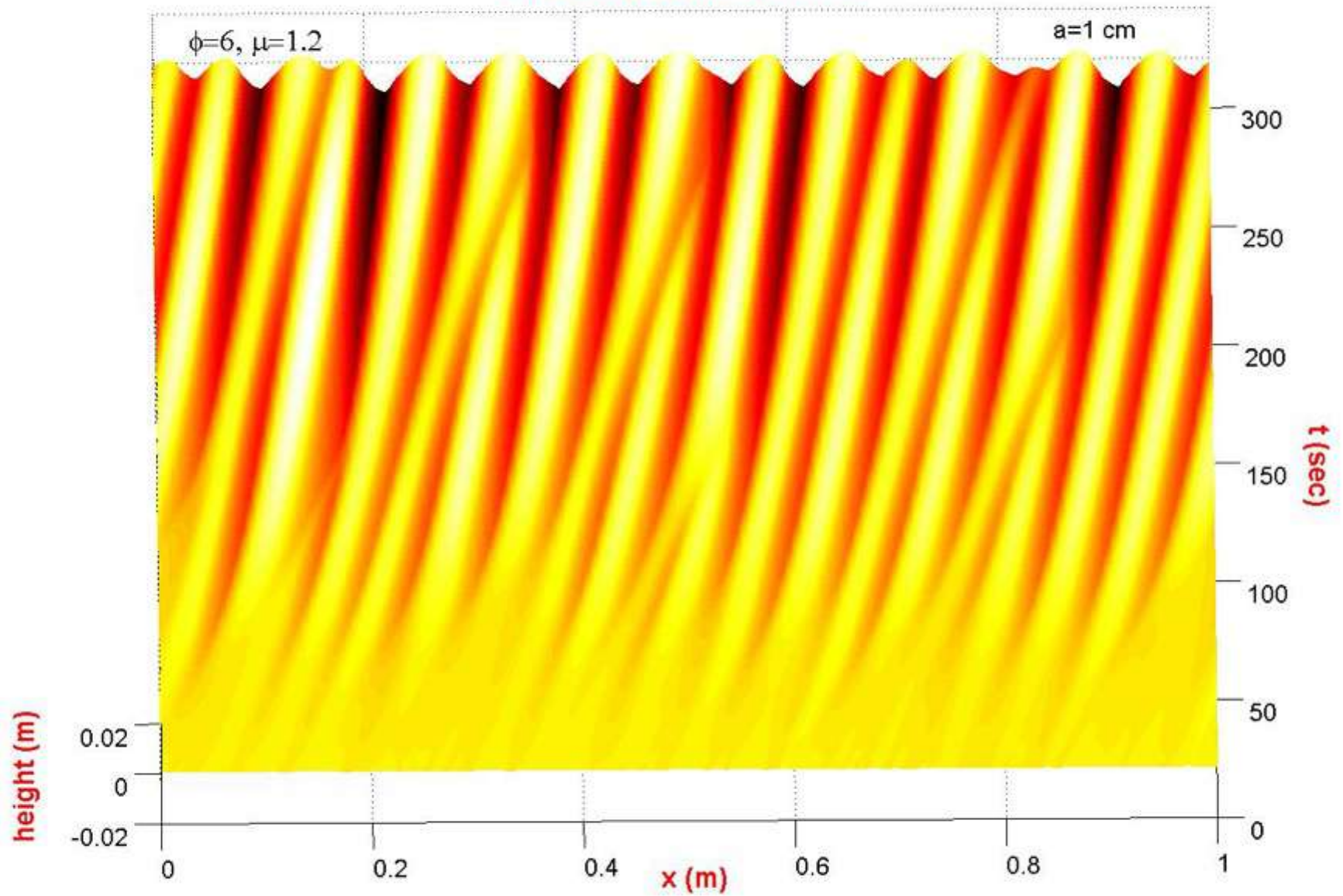
$$Q_0 = \frac{m N_r N_{im}^0 \cot \phi}{\rho(1 - n)}$$

Linear stability analysis

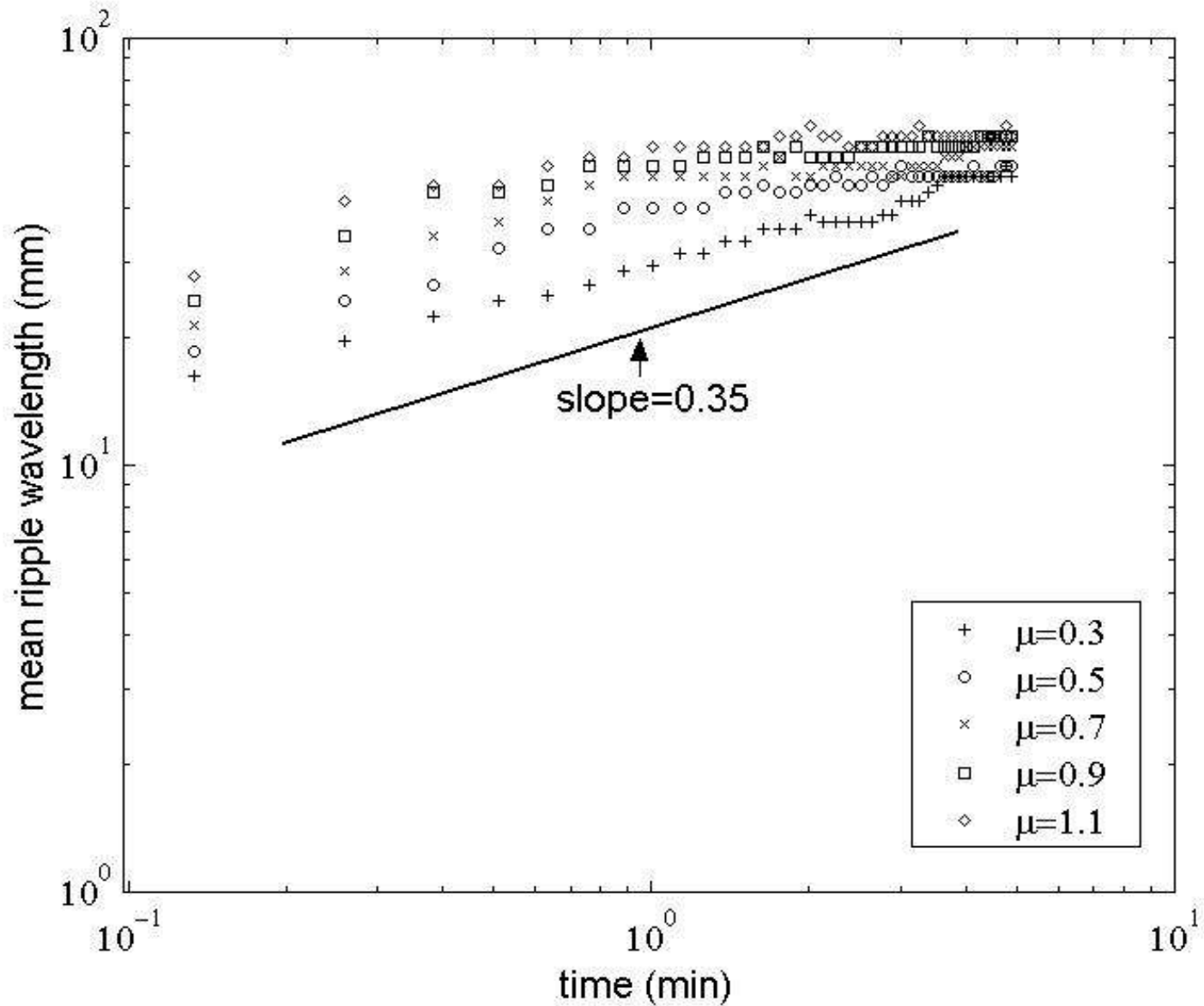
$$Q_r(x,t) = Q_r^{bare}(x,t) (1 - \mu \zeta_x)$$

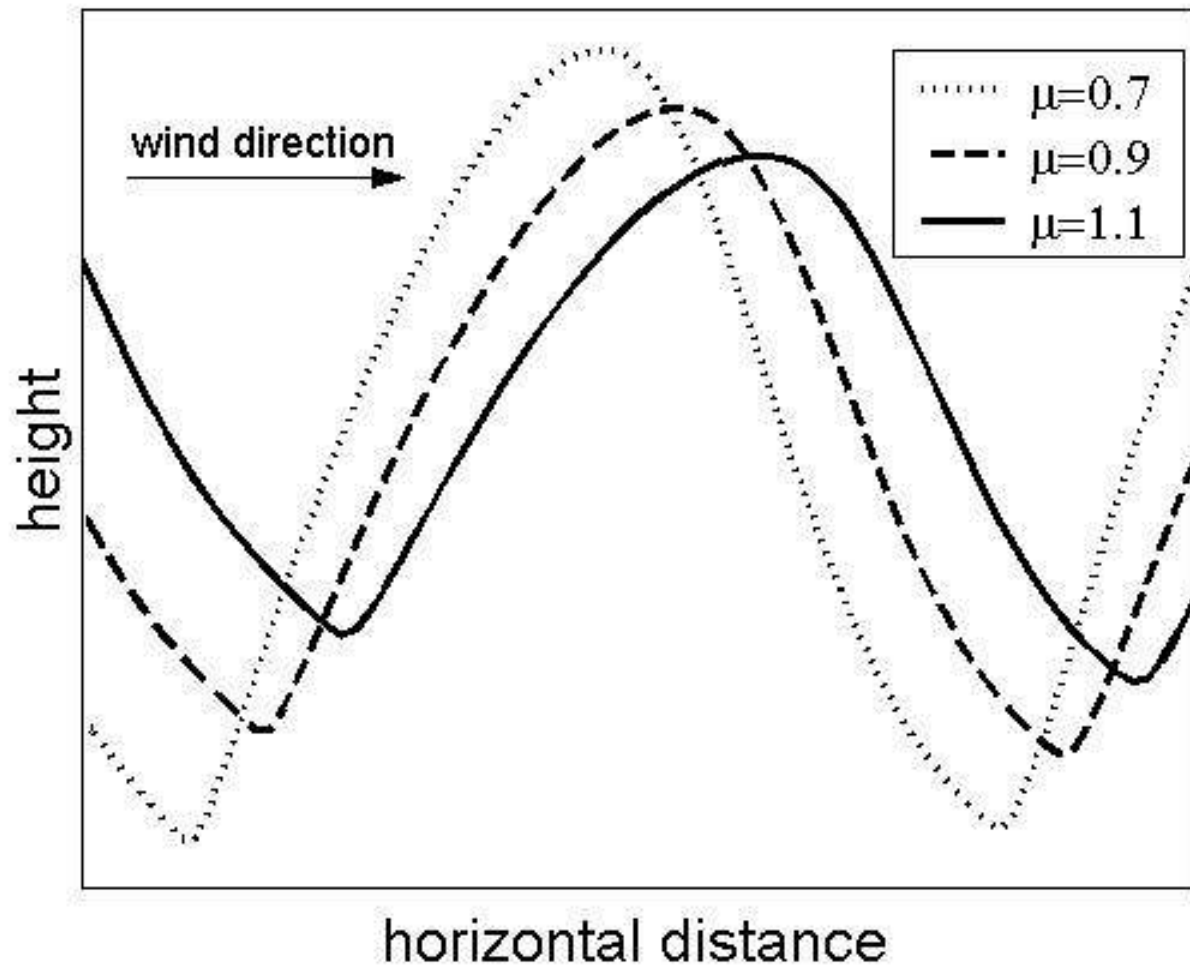


Time Evolution of Ripples



Coarsening of the ripple pattern





An extension to the 2D case:

If all the grains had
the same jump length a ,
the flux at (x,y)
in the direction ψ to the x axis
is assumed to be proportional to
the number of grains ejected
between $(x - a \cos \psi, y - a \sin \psi)$ and (x,y)

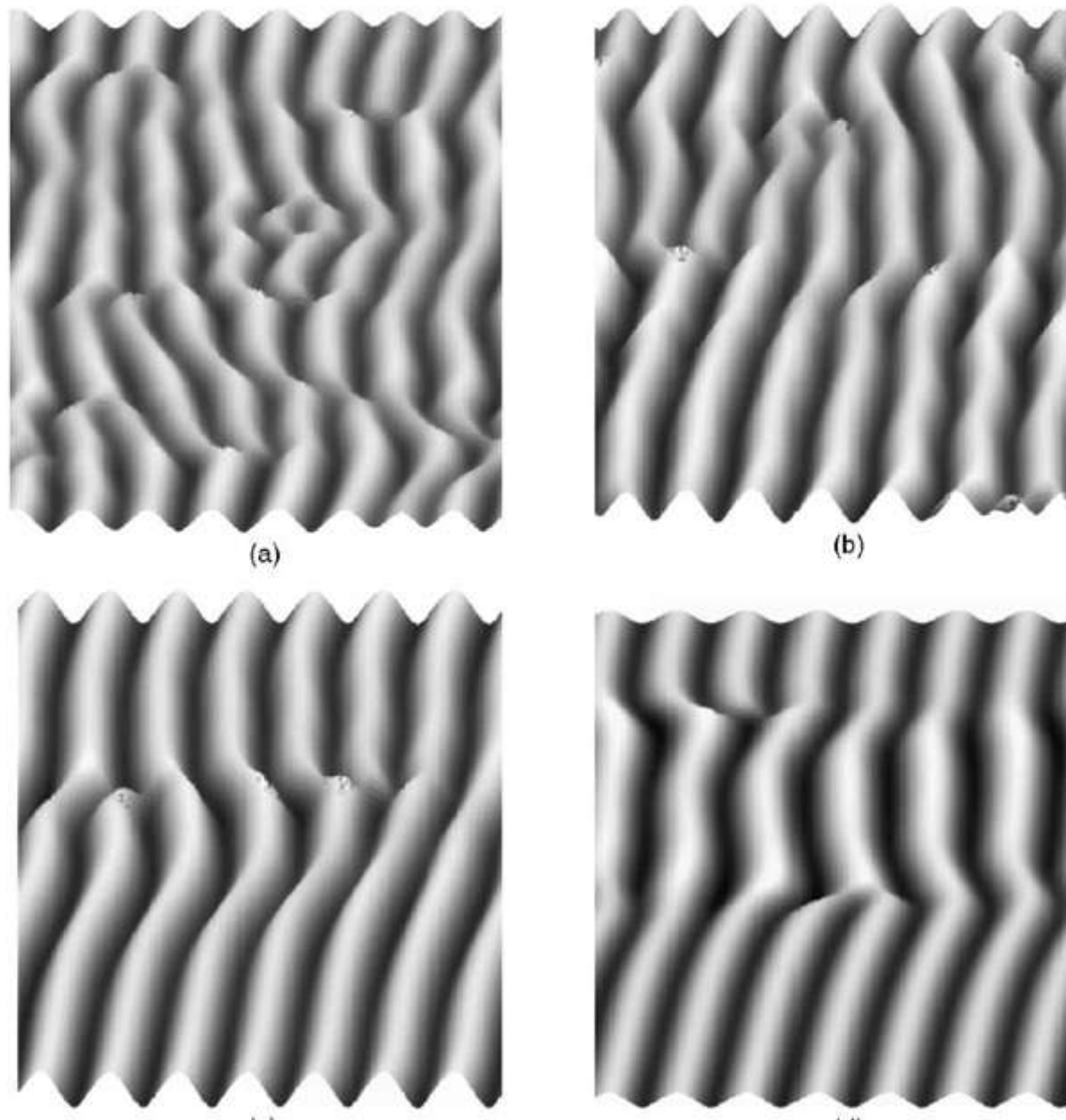
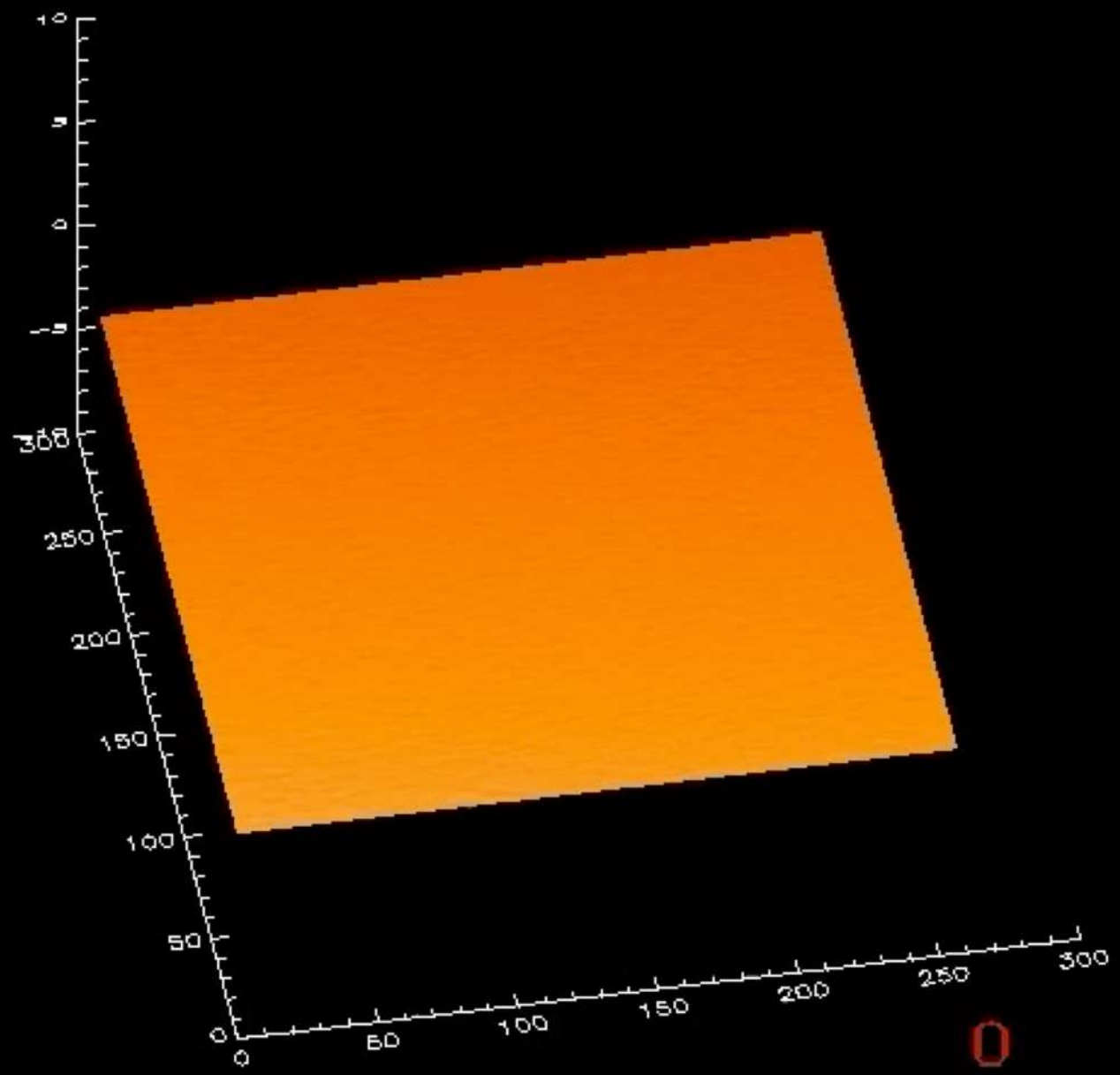
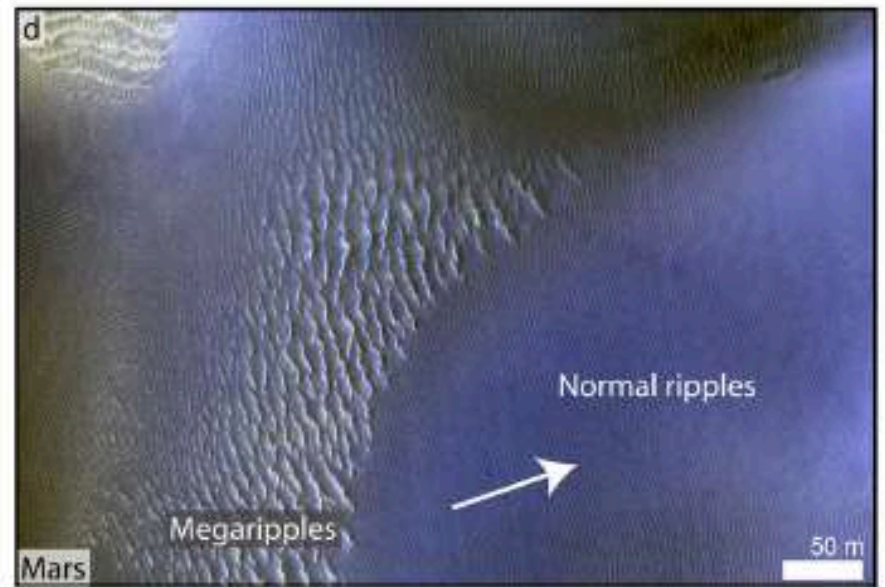
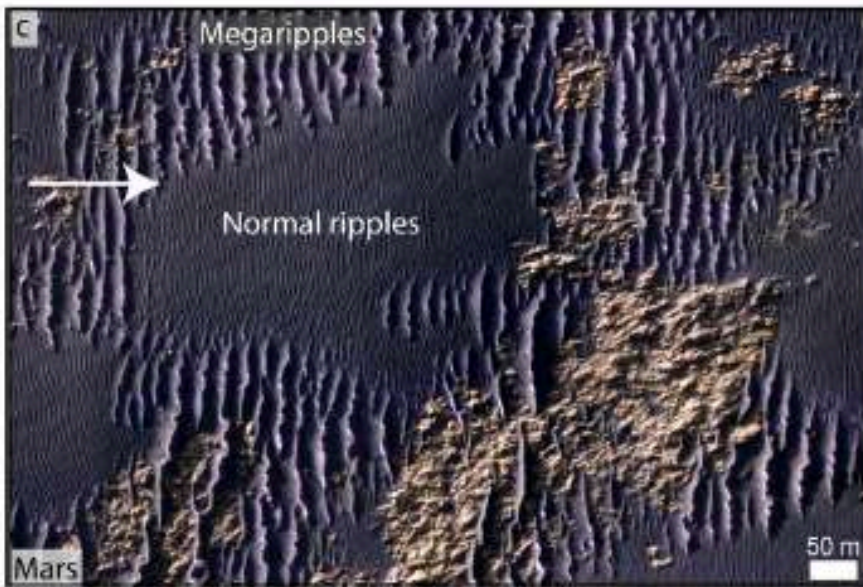


Fig. 12. Snapshots of a two-dimensional ripple field produced by the numerical integration of Eq. (28) on a 256×256 grid with periodic boundary condition, at $T = 2, 4, 8, 16$. The wind direction is from left to right. Parameter values: $\epsilon = \mu = 0.25$, $\phi = 10^\circ$, $\beta_1 = \beta_3 = 1$, $\beta_2 = \tan \phi$ and a domain size of $5\pi/2$.



0

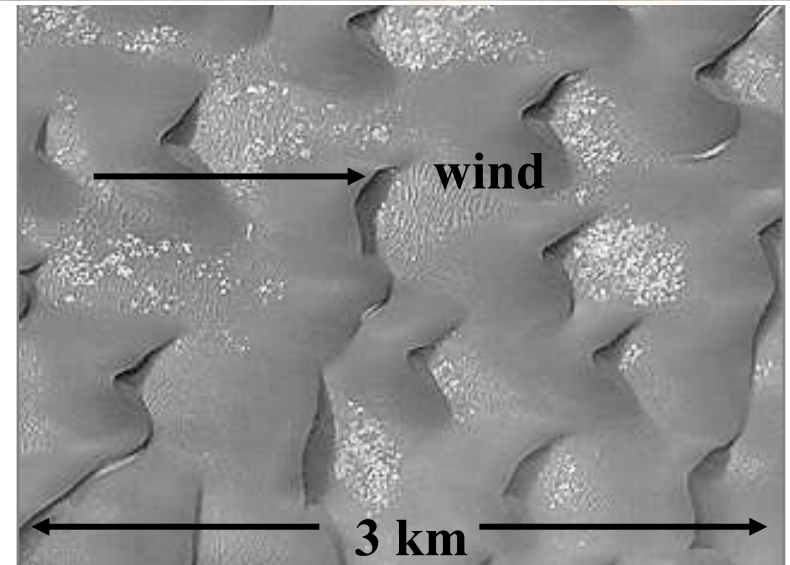


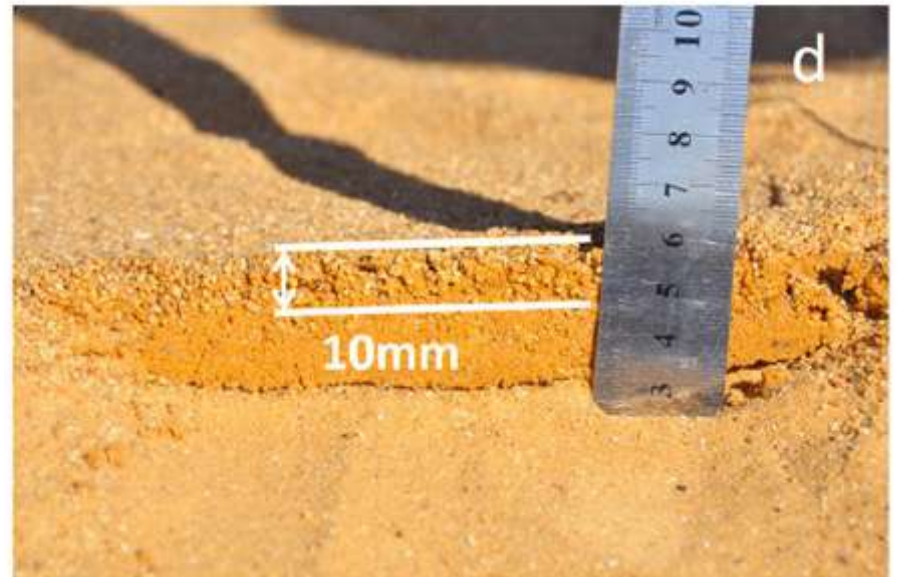
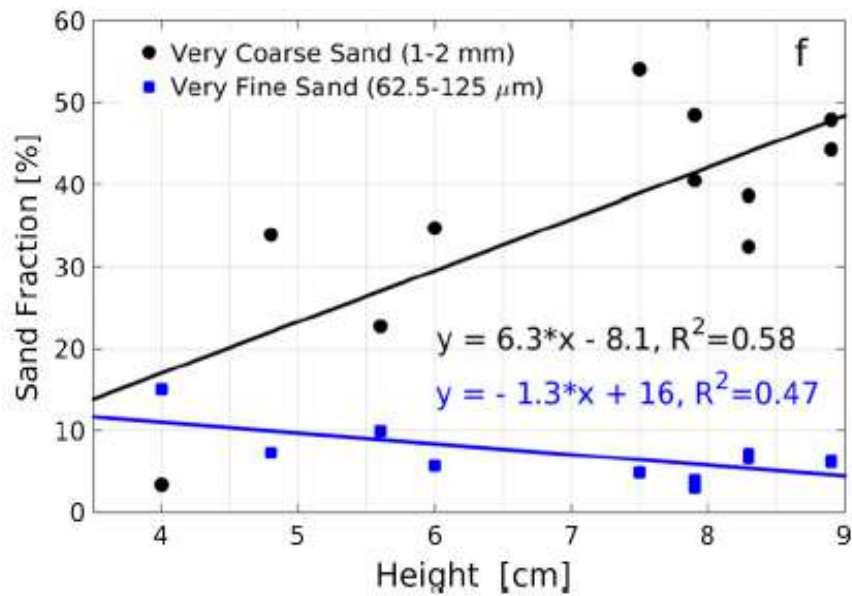
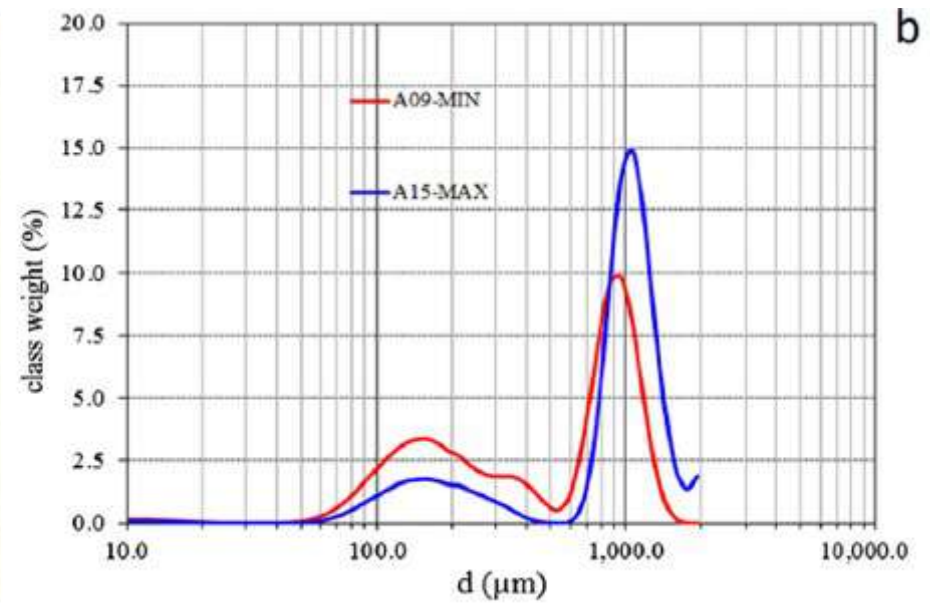
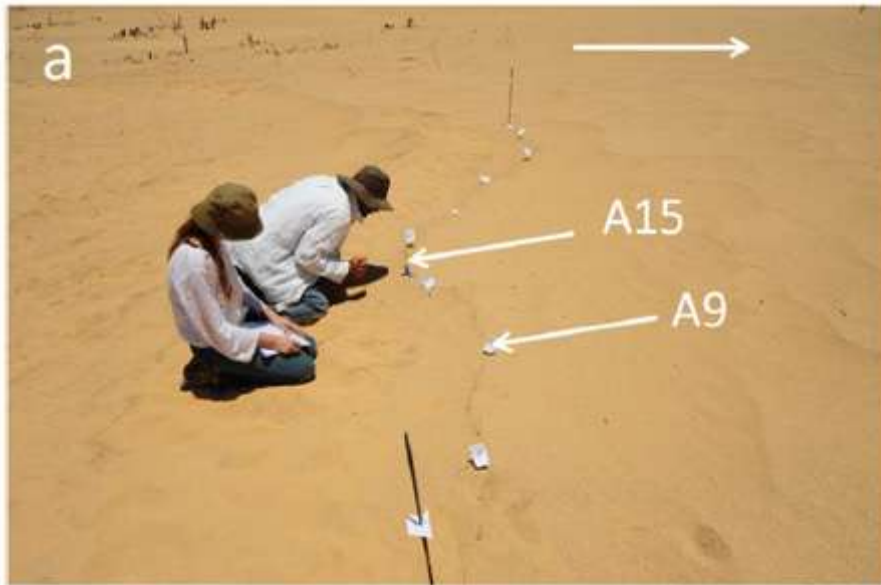
Megaripples on Earth and Mars

The origin of the transverse instability of aeolian megaripples

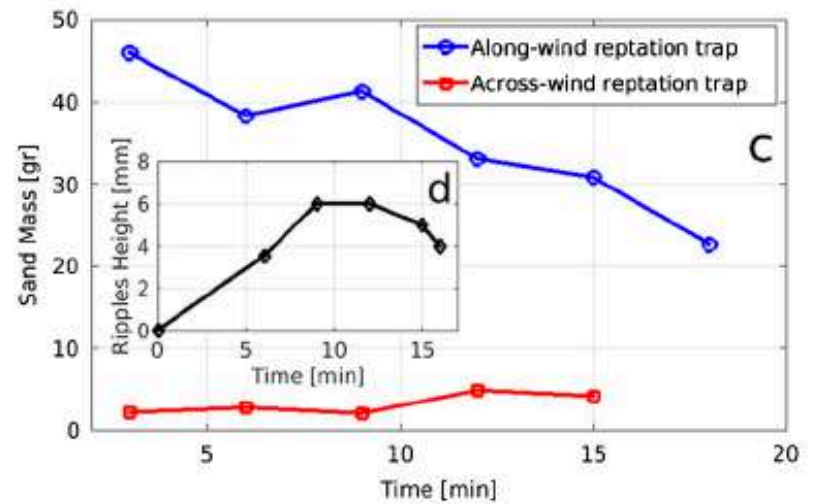
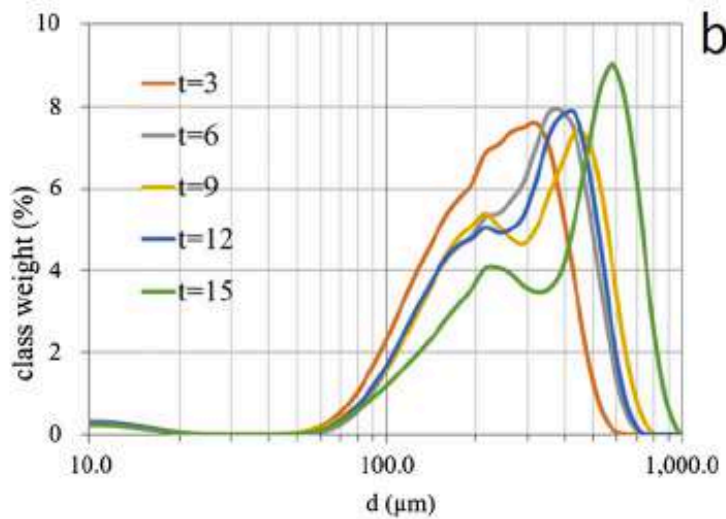
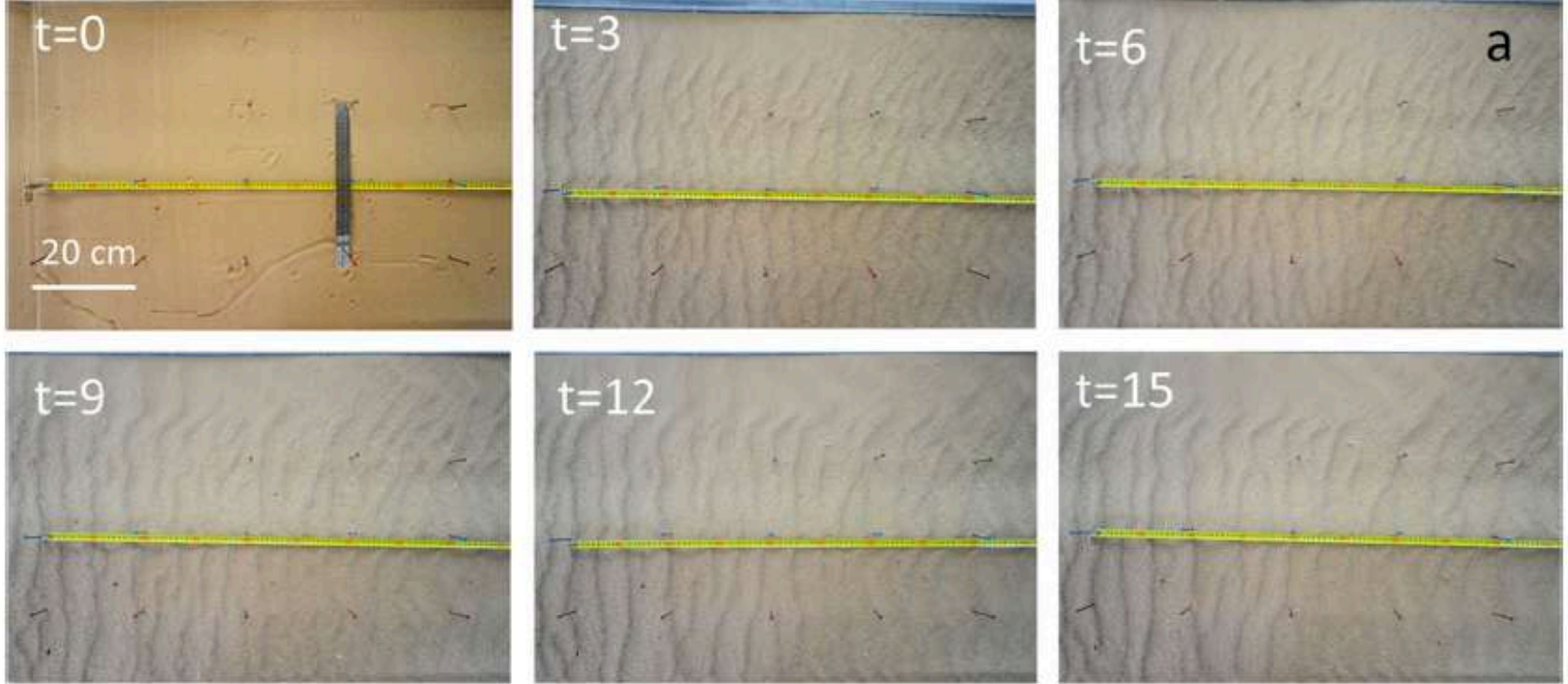
H. Yizhaq^{a,*}, G. Bel^{a,b}, S. Silvestro^{c,d}, T. Elperin^e, I.F. Kok^f, M. Cardinale^g, A. Provenzale^h,
I. Katraⁱ

Earth and Planetary Science Letters 512 (2019) 59–70

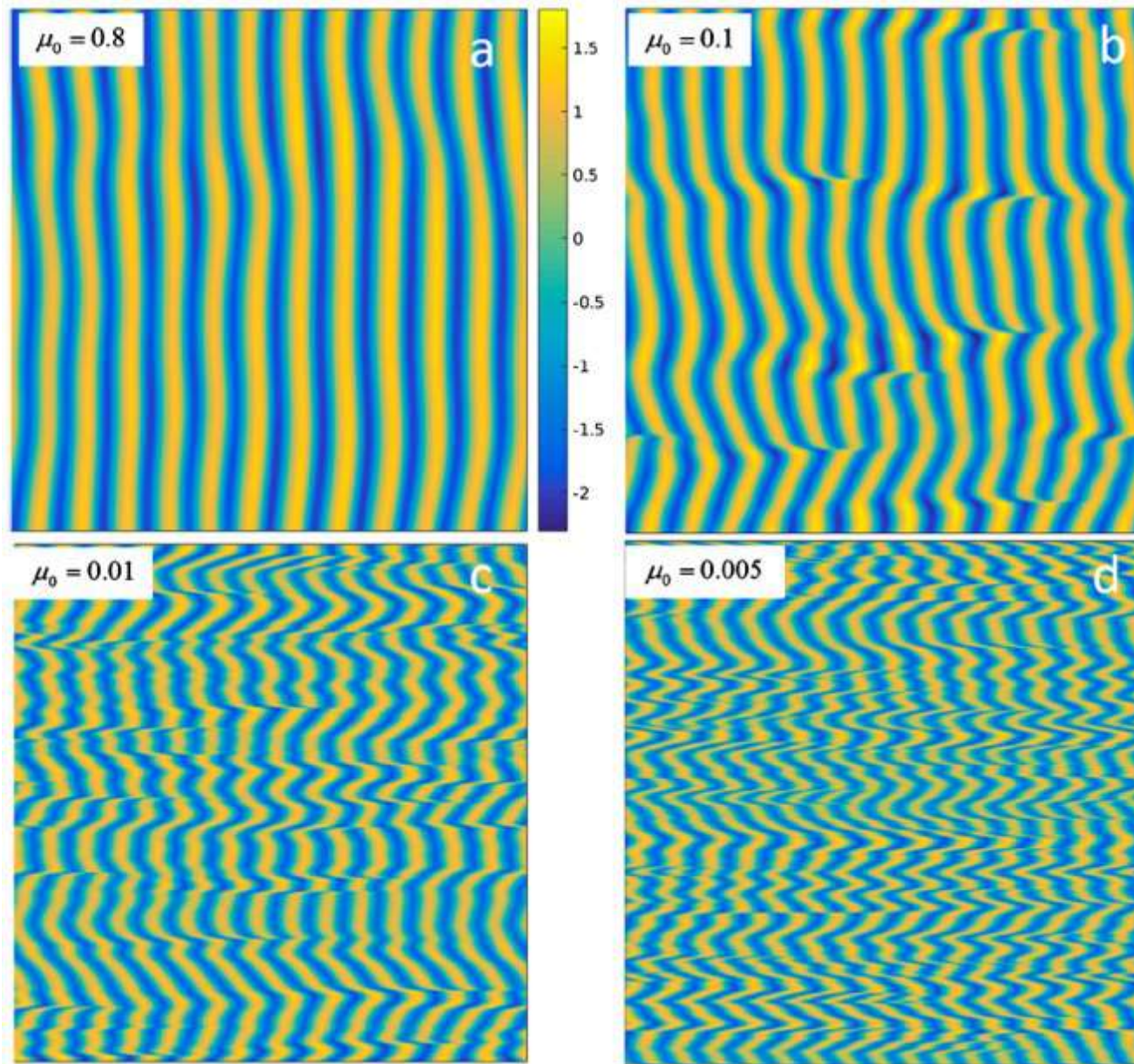




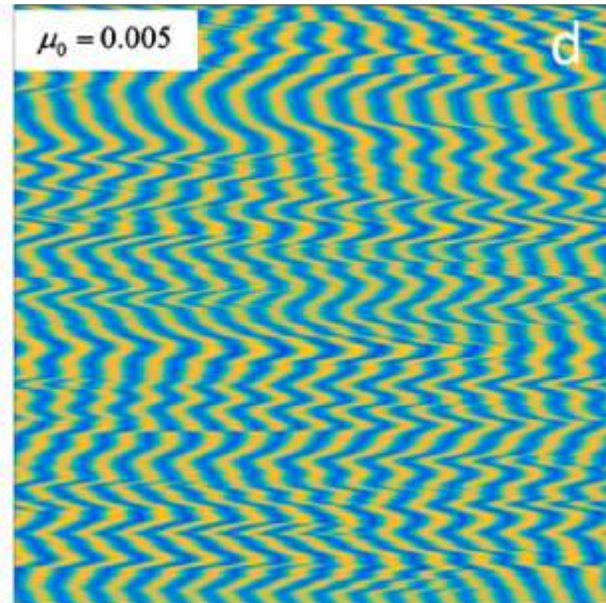
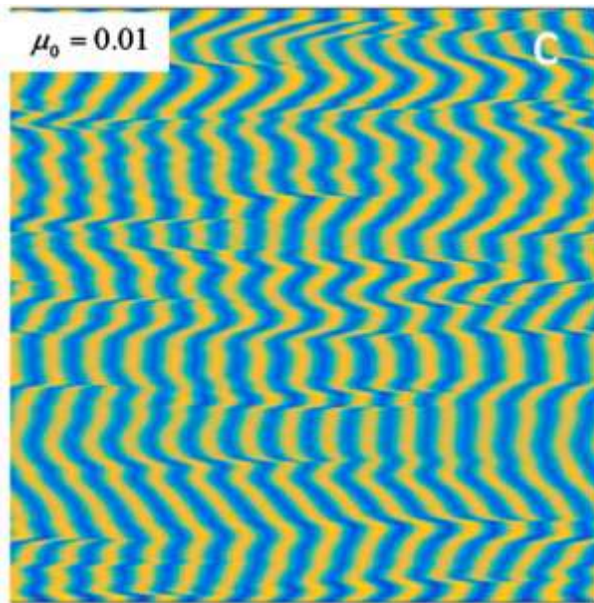
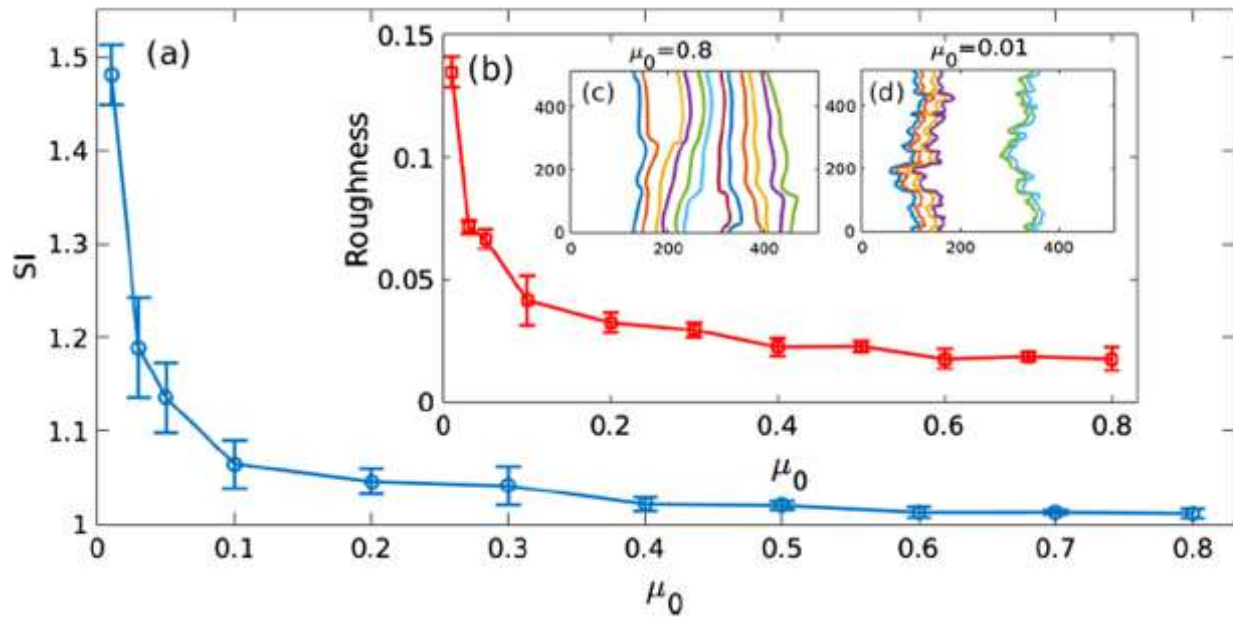
Field work



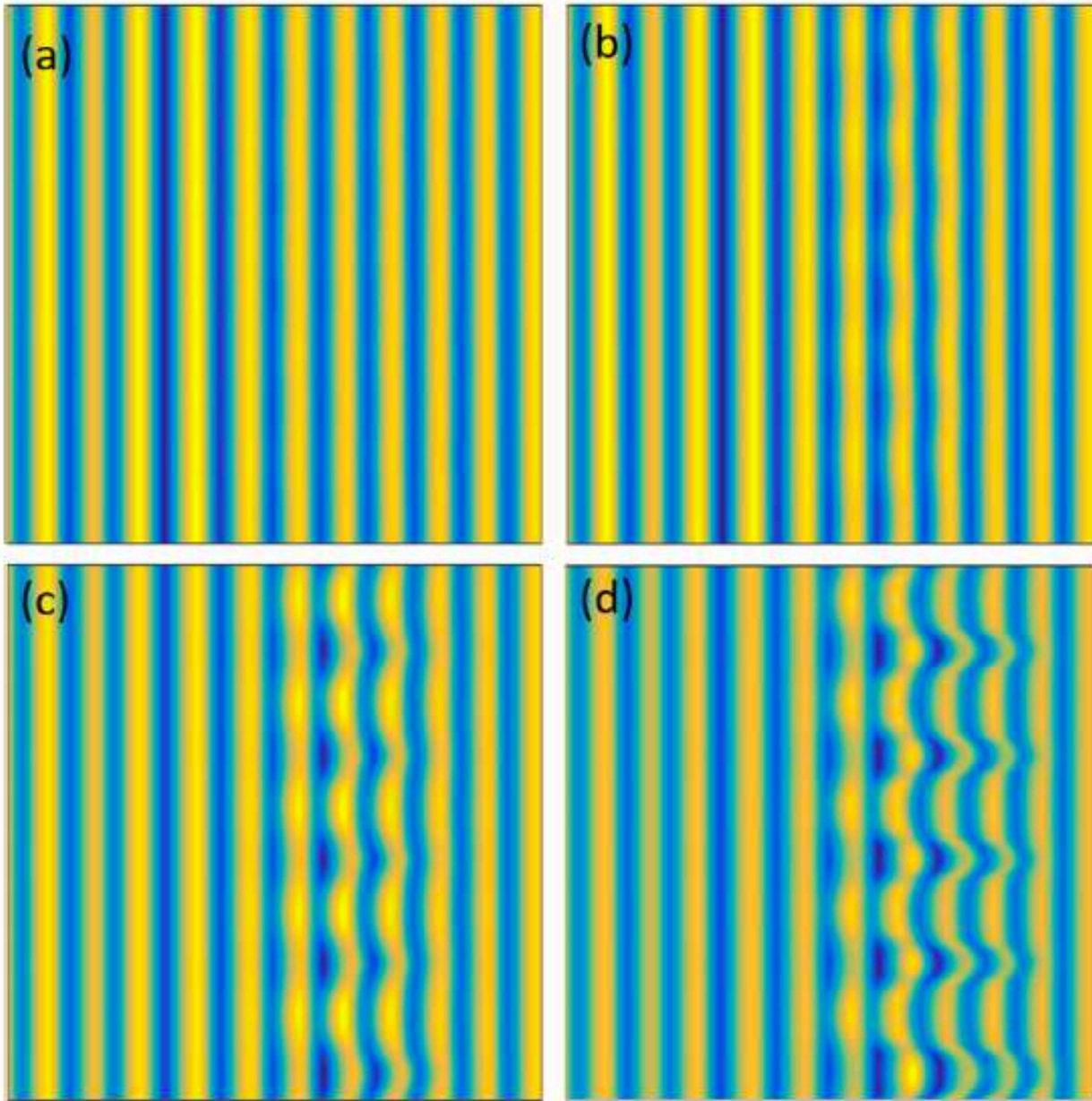
Laboratory experiments



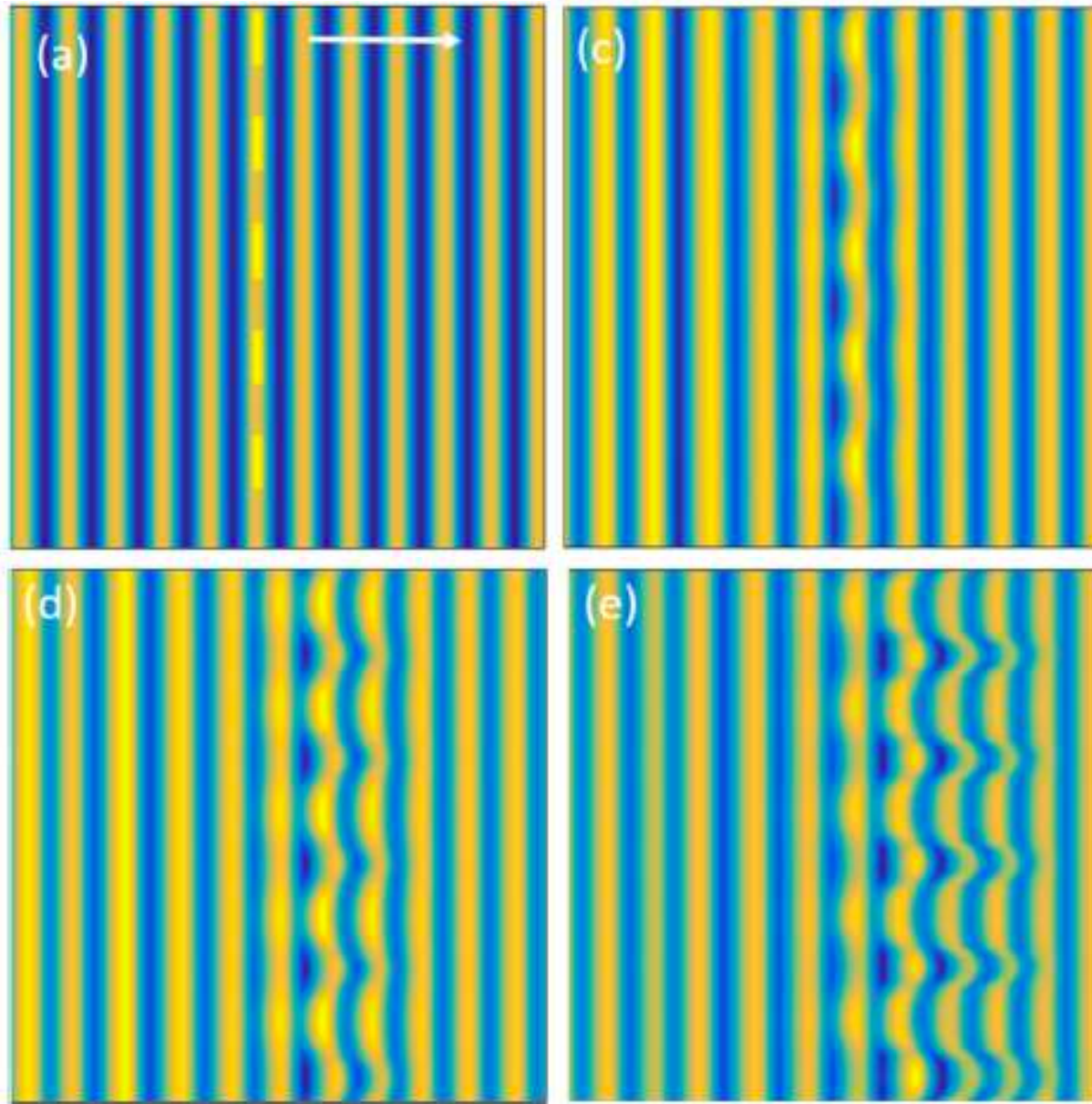
Numerical simulations



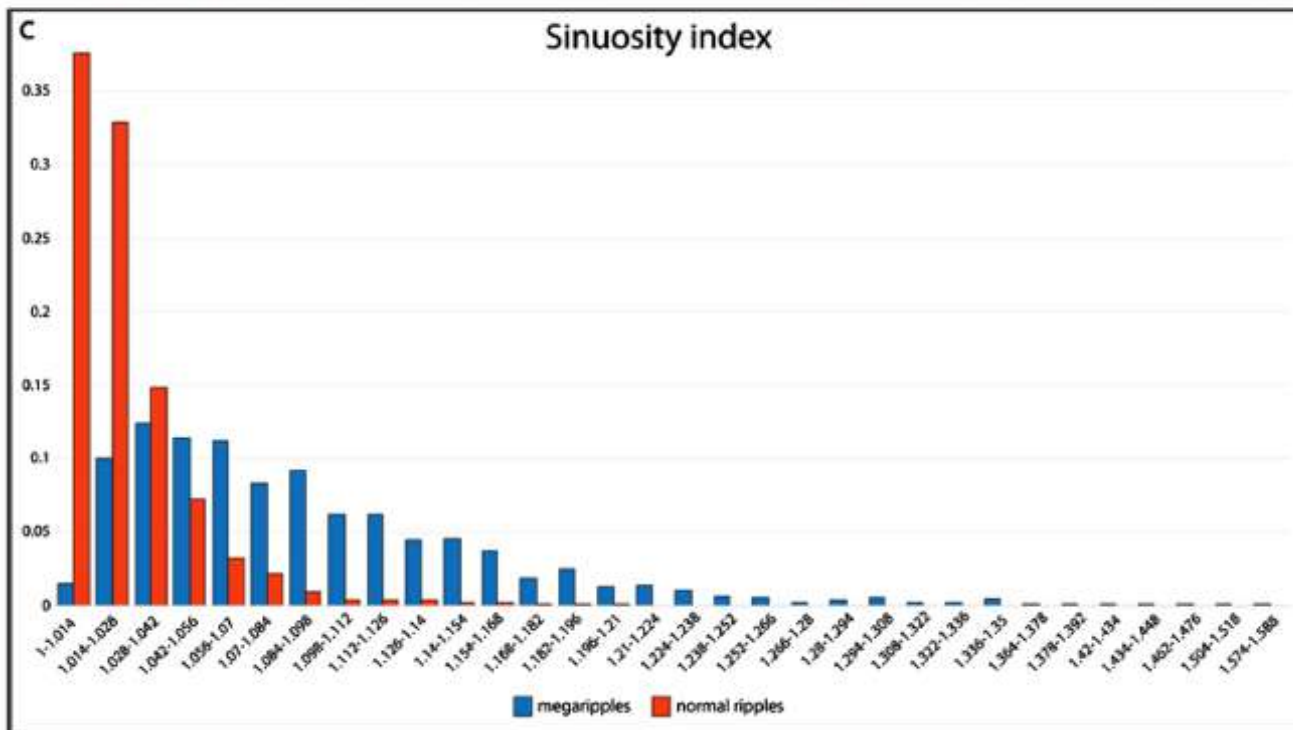
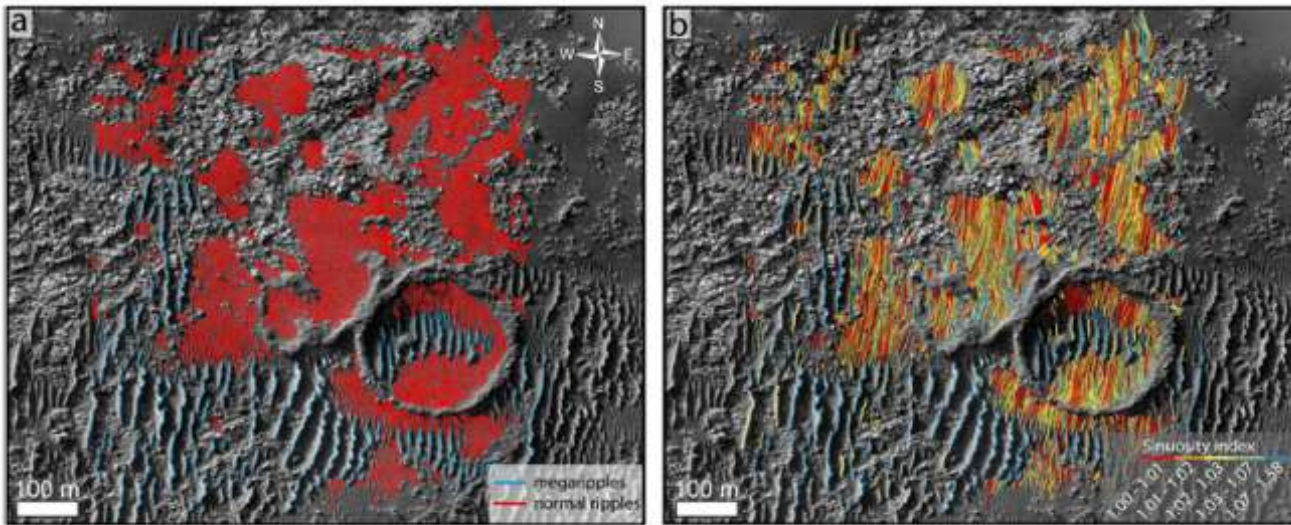
Numerical simulations



Numerical simulations



Lateral flux dependent on height



Mars case



**A much harder problem:
The dynamics of aeolian sand dunes**



**Just a beginning in the fascinating
world of (eco)geomorphological
pattern modelling**